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Abstract: Protogranular spinel-peridotite mantle xenoliths and host sodic alkaline lavas of Cretaceous to Paleogene age from $\approx 26^{\circ}\text{S}$, central eastern Paraguay and Andes are studied for: (1) trace element behaviour; (2) Sr-Nd-Pb relationships among host lavas, mantle xenoliths and hosted pyroxenes. The mantle xenoliths are distinct into two main suites, i.e. LK (relatively low in K and incompatible elements, IE) and HK (high in K and IE), both ranging from lherzolite to dunite and showing trends of "melt extraction". The hosted clinopyroxenes display extremely variable enrichment/depletion behaviours, mainly LREE. The enriched components were mostly trapped in clinopyroxenes that previously crystallized from depleted to quasi-chondritic mantle sources. Oxygen isotopes (clinopyroxene-olivine pairs) suggest equilibration temperatures higher in the HK suite than in LK suite.

Sr-Nd-Pb isotope ratios of alkaline mafic intra-plate magmatism constrain the isotopic compositions of the lithospheric mantle. Most small-volume sodic lavas were derived from a depleted lithospheric mantle source with rather uniform initial $^{143}\text{Nd}/^{144}\text{Nd}$ (0.5127-0.5128) and $^{87}\text{Sr}/^{86}\text{Sr}$ (0.7032-0.7040). The initial $^{206}\text{Pb}/^{204}\text{Pb}$ ratios are variable (18.5-19.7) at uniform $^{207}\text{Pb}/^{204}\text{Pb}$ ratios (15.60 ± 0.05). The variable Sr ,

Nd and Pb ratios are different from depleted mantle (DM), enriched (EM), or high- μ (HIMU) mantle reservoirs and probably due to radiogenic growth in a metasomatized lithospheric mantle. Notably, Sr-Nd-Pb isotope signatures of some Paraguayan xenoliths are similar to the isotopic composition of old sub-continental lithospheric mantle of the Brazilian Shield.

The isotopic Sr-Nd-Pb data and trace elements indicate that the lithospheric mantle prior to the enrichment event was dominated by a depleted component, isotopically resembling MORB source (or even more depleted?). Model ages (Nd-TDM) of clinopyroxenes and host rocks indicate tentatively main metasomatic events (fluids variously enriched in IE, REE and CO₂) at the Brasiliano cycle (i.e. 900-460 Ma). A late, likely Cretaceous metasomatism and or small degree-melting likely caused the mismatch of element ratios (Rb/Sr, Sm/Nd, U/Pb) and respective isotope systematics by redistribution of material within the lithospheric mantle, during the different phases of lithospheric thinning in the area.

Mantle Xenoliths from intracratonic Eastern Paraguay (South America Platform) and Andean Domain: a comparison

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Protogranular spinel-peridotite mantle xenoliths and host sodic alkaline lavas of Cretaceous to Paleogene age from $\approx 26^\circ\text{S}$, central eastern Paraguay and Andes are studied for: (1) trace element behaviour; (2) Sr-Nd-Pb relationships among host lavas, mantle xenoliths and hosted pyroxenes. The mantle xenoliths are distinct into two main suites, i.e. LK (relatively low in K and incompatible elements, IE) and HK (high in K and IE), both ranging from lherzolite to dunite and showing trends of "melt extraction". The hosted clinopyroxenes display extremely variable enrichment/depletion behaviours, mainly LREE. The enriched components were mostly trapped in clinopyroxenes that previously crystallized from depleted to quasi-chondritic mantle sources. Oxygen isotopes (clinopyroxene-olivine pairs) suggest equilibration temperatures higher in the HK suite than in LK suite.

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The isotopic Sr-Nd-Pb data and trace elements indicate that the lithospheric mantle prior to the enrichment event was dominated by a depleted component, isotopically resembling MORB source (or even more depleted?). Model ages (Nd-TDM) of clinopyroxenes and host rocks indicate tentatively main metasomatic events (fluids variously enriched in IE, REE and CO_2) at the Brasiliano cycle (i.e. 900–460 Ma). A late, likely Cretaceous metasomatism and or small degree-melting likely caused the mismatch of element ratios (Rb/Sr, Sm/Nd, U/Pb) and respective isotope systematics by redistribution of material within the lithospheric mantle, during the different phases of lithospheric thinning in the area.

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2

Introduction

3 Eastern Paraguay, central South America, extends in an intercratonic region including the
4 westernmost side of the Paraná-Angola-Edendeka system (PAE) and it is located between the
5 compressional Andean and extensional Atlantic systems (Fig. 1; cf. Gudmundsson & Samdridge,
6 1998). At the Paraguay latitude, the exact limit between Andean and Atlantic systems is
7 unknown due to the presence of the Chaco-Pantanal Paleogene-Neogene basin. The seismic-
8 tomography images show two high velocity features beneath Paraguay (P-waves; cf Fig. 1B), up
9 to about 200 km and more than 450 km depth, respectively, the latter probably being part of the
10 subducting Nazca slab (Liu et al., 2003). A low-velocity anomaly (Fig. 1B) in the upper mantle
11 and mantle transition zone (MTZ) was interpreted as a fossil plume by VanDecar et al. (1995).
12 However, the thinning of the MTZ has not been observed and Liu et al. (2003) suggest that
13 either this thermal anomaly does not extend into the MTZ, or, alternatively that the observed
14 anomaly is not primarily thermal, and rather dominantly compositional in origin (e.g. "veined"
15 mantle).

16 Mantle derived magmatism and mantle xenoliths of Cretaceous to Early Cenozoic age are the
17 primary sources of information on the thermal and compositional state of the uppermost mantle.
18 This contribution reviews the existing information with a focus on the evidence from mantle
19 xenoliths associated with the magmatism.

20

21

Late Mesozoic to Early Cenozoic intra-plate magmatism

22

23 Eastern Paraguay is characterized by: a) potassic alkaline-carbonatitic magmatism (ca. 140-126
24 Ma, Comin-Chiaramonti et al., 2007); Serra Geral flood tholeiites (133±1.1 Ma; Renne *et al.*,
25 1992), both represented by high-Ti and low-Ti variants (cf. Bellieni et al., 1986); c) late Early
26 Cretaceous to Paleogene sodic alkaline complexes, plugs and dykes (Velázquez et al., 2006).

Fig. 1

The most recent magmatism, is exclusive to southeast and central Paraguay, respectively (cf. Fig. 1), being concentrated at the Misiones and Asunción provinces, and characterized mainly by mafic-ultramafic sodic magmatism (Comin-Chiaramonti et al., 1986, 2001, 2007; Velázquez et al., 2002; 2006).

Pre-Cenozoic, mainly Cretaceous rift systems, developed at the Pacific side of the continent mainly on crust intensely worked in Early Paleozoic, which delimitates the western part of the Brazilian shield (Fig. 1; cf. Lucassen et al., 2002; 2007). Notably, small volumes of rock-types similar to those from Paraguay are widespread along the Central Andes (19-28°S) related to extensional structures (Lucassen et al., 2007) in the hinterland of an active continental margin (Fig. 1). The age of the magmatism spans from 119 Ma (Misiones; Velázquez et al., 2006) to 59 Ma (Asunción; Comin-Chiaramonti et al., 2007) and from about 130 Ma to 80 Ma (Central Andes; Lucassen et al., 2002, 2007). Both Paraguayan and Andean occurrences are characterized by the presence of mantle xenoliths in spinel facies, which vary in size from a few centimetres to 45 cm (Comin-Chiaramonti et al., 2001; Lucassen et al., 2005). Notably, assuming a diameter of 45 cm, corresponding to the size of the largest xenoliths, a density of 3.3 g/cm³ and an origin of the host liquids from depth ~70-75 km (~boundary between garnet and spinel peridotite), the transport of the xenoliths to the surface occurred in a very short time (e.g. less than 9 days: s. Spera, 1984; cf. Comin-Chiaramonti et al., 1991).

Table 1.

Ultramafic xenoliths, occasionally hosted by alkaline lavas and generally associated with the magmatism in continental rift areas, provide the opportunity for a direct sampling of the subcontinental mantle. Several studies on mantle xenoliths have pointed out that incompatible

53 trace elements, radiogenic isotopes and the major element chemistry are largely decoupled from
54 the coherent variation that would be expected by simple extraction during partial melting of a
55 peridotite at mantle conditions (cf. Frey and Prinz, 1978). In particular, variable enrichment of
56 incompatible elements is commonly observed in contradiction to the refractory nature of the
57 xenoliths, thus requiring additional processes besides partial melting and fractionation. Such
58 processes have been generally called *metasomatic* (e.g. Menzies and Hawkesworth, 1987).
59 Following the latter authors, metasomatism of residual peridotites by volatile-charged fluids or
60 small-volume melts carrying incompatible elements cause different styles of enrichments.
61 Metasomatically changed parent/daughter element ratios could develop isotopic
62 inhomogeneities by radioactive insitu decay in a non-convective, lithospheric
63 mantle. Geochemical and isotopic data of mantle xenoliths provide valuable information about
64 the nature and timing of differentiation and possible metasomatism of mantle domains.

66 **Characteristics of the host lavas**

67
68 The Asunción and Misiones lavas ($^{40}\text{Ar}/^{39}\text{Ar}$ ages 58.4 ± 2.1 and 118.3 ± 0.9 , respectively; Comin-
69 Chiamonti, et al., 2007) are melanephelinites and subordinate ankaratrites, according to the
70 nomenclature after De La Roche (1980, 1986; cf. Fig. 2) and Le Maitre (1989), with mg#
71 $[\text{MgO}/(\text{MgO}+\text{FeO})]$, assuming $\text{Fe}_2\text{O}_3/\text{FeO}$ ratio = 0.21] ranging from 0.64 to 0.67 (Comin-
72 Chiamonti et al., 1991). Representative and average analyses are reported in Table 1.

73 The Andean lavas and dykes are ankaratrites (Finca del Rodeo, mg# 0.67), basanites (Las
74 Conchas + Cadillal and Betanzos, mg# 0.70 and 0.59, respectively) and tephrites (Belén, mg
75 # 0.57). K-Ar data (Lucassen et al., 2007) give 82 ± 4 Ma (Betanzos), 96 ± 4 Ma (Finca del
76 Rodeo), and about 100 to 130 Ma (Las Conchas + Cadillal and Belén, respectively).

77 All these rock-types belong to the sodic alkaline suites (inset of Fig. 2), having $\text{Na}_2\text{O}/\text{K}_2\text{O}$ ratio
78 ranging from 2.12 to 5.85. Their phenocryst population is characterized by olivine (1-7 vol%,

79 Fo 89-85; 1-6 vol% microphenocrysts, Fo 82-77 mole%), clinopyroxene (1-6 vol%; mg# ~0.8),
80 Ti-magnetite (0.3-0.7 vol%; 37.9 ulv. mole%), and occasional phlogopite, e.g., in the Betanzos
81 rocks. The hypocrySTALLINE groundmass is made up by clinopyroxene (39-46% vol%, mg#
82 ~0.75); olivine (3-6 vol%, Fo 74-76 mole%); magnetite (4-7 vol%, 42.8% ulv.mole%);
83 nepheline (16-21 vol%); and glass (11-25 vol%).

84

85 **Fig. 2.**

86

87 Trace element and REE patterns display a similar topology (Fig. 3). The major difference is the
88 variable Rb contents, which is higher in the Paraguayan rocks than in their Andean analogues
89 (K/Rb: 130.2 to 205.4 in Paraguay vs. 342-1142 in the Andes). The overlapping compositions
90 of Paraguayan and Andean magmatism are apparent, except for the Las Conchas+Cadillal
91 basanites (mg# 0.70) that are Rb and LREE depleted with respect to the other rocks. K is
92 systematically strongly depleted with respect to neighbouring elements, whereas (Ta)-Nb and Sr
93 tend to show positive spikes. (Eu/Eu*) is negative in Paraguay, ranging from 0.75-0.85 and near
94 the unity in the Andean domain (1.01-1.03; cf. Tab. 1 and Fig. 3B).

95 Melting models indicate that these rocks derived from liquids representing 4-6% degrees of a
96 garnet-peridotite (cf. Comin-Chiaramonti et al., 1997) and that these liquids cannot be derived
97 from sources similar to those represented by the enclosed spinel-peridotite xenoliths.

98

99 *Radiogenic isotopes*

100

101 Averaged Sr-Nd-Pb isotopic compositions (time integrated) representative of the Cretaceous to
102 Paleogene magmatic rocks are reported in Table 2.

103 Due to the high Sr and Nd contents of the mafic alkaline rocks compared to crustal values of
104 these elements, it is reasonable to assume that Sr-Nd isotopic compositions have not been

105 significantly affected by crustal contamination processes (see below). Therefore, Sr_i and Nd_i
106 isotopic compositions are considered to represent the isotopic composition of their mantle
107 source (cf. Comin-Chiaramonti et al., 1997; Lucassen et al., 2007).

108 The Sr-Nd isotopic variations (Fig. 3C-D) could be explained: a) by distinct portions of a large-
109 and small-scale heterogeneous lithospheric mantle source, where the small-scale heterogeneity
110 is required to explain the variations in the Sr_i and Nd_i ratios of each magmatic event; or b) in
111 terms of mixing by involving an enriched mantle component with (EMI or K-ASU end-member
112 for subcontinental lithospheric mantle?; cf. Comin-Chiaramonti et al., 1997; Antonini et al.,
113 2005) and depleted mantle component(s) like DMM and/or HIMU. Depleted mantle sources
114 appear to be important in the Argentina Central Rift (Lucassen *et al.*, 2002) and also in the
115 Ayopaya alkaline-carbonatitic province, Bolivian Central Andes, which represents the
116 petrologically most exotic expression of large-scale Mesozoic continental rifting along the
117 western South American continent (Schultz *et al.*, 2004). In any case (Fig. 3C-D), it is inferred
118 that depleted mantle domain(s) played progressively a major role in the genesis of alkaline
119 magmatism in Eastern Paraguay and Argentina since Cretaceous up to Paleogene times (Comin-
120 Chiaramonti et al., 1999; Lucassen et al., 2002, 2007).

121

122

123 **Fig. 3.**

124

125 Pb isotopes are believed to discriminate between EM, DMM, and HIMU components (cf.
126 "Mantle components" reviews e.g. Zindler and Hart, 1986; Hofmann 1997). The initial Pb
127 isotopic compositions of the Paraguayan K-magmatism, most "primary" rock-types (cf.
128 Antonini *et al.*, 2005), have initial $^{206}Pb/^{204}Pb$, $^{207}Pb/^{204}Pb$ and $^{208}Pb/^{204}Pb=16.888-17.702$,
129 15.433-15.620 and 37.156-37.915, respectively. The initial lead isotopic ratios of the tholeiitic
130 Paraguayan rocks generally agree with the Brazilian analogues reported by Marques et al.

131 (1999). On the other hand, the sodic alkaline rocks have different Pb isotopic compositions and
132 differ from the potassic analogues, having initial compositions $^{206}\text{Pb}/^{204}\text{Pb}=18.31-19.39$,
133 $^{207}\text{Pb}/^{204}\text{Pb}=15.56-15.67$ and $^{208}\text{Pb}/^{204}\text{Pb}=37.97-39.16$.

134 The role of possible distinct mantle components in the genesis of the Cretaceous to Paleogene
135 magmatism are summarized in Fig. 3C-E. In general, the potassic rocks follow the same trend
136 as the Early Cretaceous tholeiites and the Cretaceous and Paleogene sodic rocks. It is interesting
137 to observe that the isotopic compositions trend towards the HIMU field rather than to the DMM
138 one, particularly in the $^{207}\text{Pb}/^{204}\text{Pb}$ vs. $^{87}\text{Sr}/^{86}\text{Sr}$ and $^{143}\text{Nd}/^{144}\text{Nd}$ diagrams, suggesting that a
139 HIMU component played an important role in the whole Cretaceous to Paleogene magmatism.
140 However, the Pb isotopic compositions of some alkaline rocks (e.g. potassic rocks with highest
141 $^{207}\text{Pb}/^{204}\text{Pb}$) tend towards the field of the basement rocks and, therefore, some contribution of
142 crustal components can not be ruled out.

143 Alternatively, the observed trends can be modelled by mixing process between the K-ASU
144 magmas (as defined by Comin-Chiaramonti et al., 1997) with the lowest initial $^{206}\text{Pb}/^{204}\text{Pb}$ ratios
145 and St. Helena HIMU magma-types. The calculated patterns show that a) 10-30% of the St.
146 Helena (HIMU end-member) is consistent with the isotopic heterogeneity observed in the
147 potassic magmatism; b) the HIMU component appears to be dominant in the generation of the
148 sodic magmatism.

149

150 **Table 2.**

151

152 **Mantle xenoliths**

153

154 The mantle xenoliths are mainly spinel-lherzolites, harzburgites and subordinate dunites and
155 correspond to the Cr-diopside series of Wilshire and Shervais (1975). The dominant texture is
156 protogranular, rarely tabular or porphyroclastic (Demarchi et al., 1988; Lucassen et al., 2005).

157 The Paraguay mantle xenoliths contain variable amounts of glassy patches ("blebs"), and glassy
158 drops in clinopyroxenes. The latter show an overprinted spongy texture, characteristic of the
159 clinopyroxenes of H-K suite (see below). The blebs (mg# 0.88-0.91) mainly consist of a glassy
160 matrix containing microlites of olivine (mg# 0.91-0.92), clinopyroxene (mg# 0.91-0.93), Cr-
161 spinel and (rarely) phlogopite (mg# 0.86-0.92). They are considered to have been formed by
162 decompression melting of amphibole (cf. Wang et al., 2007) and phlogopite (Comin-
163 Chiamonti et al., 1986; 2001). In the Andes, clinopyroxene in some samples shows rims of
164 secondary pyroxene with sieve texture and pockets of devitrified glass (Lucassen et al., 2005). ,
165 Blebs and secondary pyroxene rims are likely caused by decompression melting during the
166 rapid ascent. The rapid ascent is also testified (in addition to petrophysical evidences and the
167 persistence of glassy drops and patches mentioned above) by the lack of any apparent
168 interactions between the mantle xenoliths and host lavas (Demarchi et al., 1988) and by lack of
169 phase transitions in the associated crustal xenoliths (e.g. plagioclase-microcline, quartz-
170 sillimanite pairs etc for which the calculated T-P equilibria are 570-700°C and 4-8.5 kbar,
171 respectively; cf. Orú et al., 1990).

172 The Paraguay xenoliths are characterized by a large range of K₂O contents (0.02 to 0.51 wt %).
173 Some xenoliths have K₂O abundances comparable or even higher than those reported for
174 metasomatized mantle peridotites (e.g. Frey and Prinz, 1978; Roden et al., 1984), resembling in
175 some case amphibole-mica-apatite bearing mantle-xenolith suites (cf. 'O Reilly and Griffin,
176 1988). K₂O contents and the occurrence of blebs and glassy drops allow to group the Paraguay
177 xenoliths into two main suites, i.e. a low-K suite (LK, K₂O < 0.1 wt %), and a high-K suite with
178 abundant glassy drops and/or variable amounts of blebs (HK, K₂O ≥ 0.2 wt %). On the other
179 hand, the Andean xenoliths have K₂O content < 0.08 wt% (average 0.02± =0.02 wt%)
180 Complete sets of chemical analyses are in Demarchi et al (1988) and Lucassen et al. (2005).
181 Coherent variations of major element contents in both suites follow a dunite-lherzolite sequence
182 trending to primitive mantle compositions (Fig. 4A). The population is mainly represented by

183 harzburgitic compositions, being mostly within 0.55-0.63 range of $(\text{SiO}_2 + \text{Al}_2\text{O}_3)/(\text{MgO} + \text{FeO}_T)$
184 molar ratio. The residual character of the mantle xenoliths, believed to be consistent with
185 melting and basalt-component removal, is shown also by the decrease in the cpx/opx modal
186 ratio with decreasing modal cpx, which fits the model variation trend induced by partial melting
187 of lherzolite (Fig. 4B).

188 **Fig. 4.**

189 *Mineral Phases*

190 *Clinopyroxene* (Cpx, mg# 0.90-0.96; 1 to 26 vol%) exhibits a variable morphology and
191 composition in the Paraguayan samples, ranging from well preserved (lherzolite) to relics-
192 spongy (dunite) crystals. The most common Cpx habitus, occurring in the cores and margins of
193 the xenoliths of the HK suite, is represented by crystals with partially to completely "spongy"
194 texture, due to glassy drops. Usually Cpx shows lamellar spinel (rarely orthopyroxene)
195 exsolutions cut by the glassy drops. Because of the spinel exsolutions and glassy drops, Cpx has
196 been analyzed (microprobe analyses) both on spots and "traverse" (the latter obtained by shifting
197 the sample under defocussed beam in order to provide an estimate of composition before
198 unmixing). The "traverse" compositions of spongy Cpx from highest K_2O -bearing xenoliths are
199 anomalously rich in K_2O (up to 1.22 wt %). This reflect the occurrence of the glassy drops (up ~
200 7.5% of the whole cpx composition, as from mass balance) whose compositions are variable,
201 but are systematically rich in K_2O , Al_2O_3 , Rb, Sr, and Ba (cf Demarchi et al. 1988).
202 Clinopyroxenes in the Andean samples are of rather restricted composition and comprise
203 (aluminian, sodic, \pm chromian) augite to diopside with quadrilateral components > 87%
204 (Lucassen et al., 2005)
205 *Orthopyroxene* (Opx, mg# 0.90-0.93; 3 to 30 vol%) is unzoned and rarely shows spinel
206 exsolutions.

207 *Olivine* (Ol, 60 to 97 vol%) is chemically homogeneous within each samples and varies in
208 composition mainly from Fo 89 (Iherzolites) to Fo 93 (dunites) mole%.

209 *Spinel* (Sp, 0.6 to 4 vol%) from xenoliths with $K_2O \geq 0.20$ wt % (see below) is generally
210 characterized by Cr increase from the core to the rim, particularly grains inside the blebs, while
211 spinels from xenoliths with $K_2O < 0.1$ wt % are unzoned (Demarchi et al. 1988). The Cr/(Cr+Al)
212 atomic ratio shows a main variation range between 0.10 and 0.70.

213 *Sulfides* (pyrrhotite and pentlandite) and *apatite* occur in very small amounts as dispersed grains
214 in almost all the xenoliths.

215 *Carbonate*

216 In Paraguay carbonates are present in pockets in the glassy patches, or as carbonate veins along
217 the grain boundaries. They are Mg-rich calcite ($MgCO_3$ about 5-10 wt% and < 0.6 wt%).
218 Measured $\delta^{18}O\%$ (SMOW) and $\delta^{13}C\%$ (PDB) in two representative samples (HK suite) give
219 $+8.5 \pm 0.5$ and -7.4 ± 0.9 , respectively, i.e. values within the field believed typical of primary
220 carbonatites, according to Keller and Hoefs (1995).

221 In the Andes occurrences, Lucassen et al. (2007) report the presence of carbonatized peridotites,
222 similar in shape and size to the other peridotite xenoliths. Minor clinopyroxene and spinel and
223 occasional phogopite are set in a matrix of carbonate grains. $\delta^{18}O\%$ and $\delta^{13}C\%$ range from
224 $+17.3$ to $+20.7$ and from -5.3 to -8.9 , respectively, and indicate post magmatic and/or deuteric-
225 groundwater processes in the carbonate formation (cf. Comin-Chiaramonti et al., 2005).

226 The "residual" character of the xenolith suites indicated by their major element compositions is
227 also reflected by the compositional variations of mineral phases (Demarchi et al., 1988; Petrini
228 et al., 1994), i.e. Mg/ Fe_{tot} (atomic ratio) of the orthopyroxene (9.7 to 11.8) is correlated with
229 the Fo content of the olivine (90 to 93) and with Mg/ Fe_{tot} of clinopyroxene (9.2 to 15.3), as
230 well as with the Cr/(Cr+Al) ratio of the spinel (0.10 to 0.70).

231

232 *Thermobarometry*

233

234 In Paraguay, intercrystalline equilibration temperatures for orthopyroxene-clinopyroxene pairs
235 (Wells geothermometer, 1977) and for olivine-spinel pairs (Fabriès geothermometer, 1979) vary
236 between 862 and 1075 °C and between 748 and 968°C, respectively. Intracrystalline pressures
237 (calculated according to Mercier, 1980) vary between 11-23 Kb and 17-20 Kb for clinopyroxene
238 and orthopyroxene, respectively (cf. Petrini et al., 1994) in both the two suites, where the higher
239 pressures were obtained for the more depleted xenoliths of the LK-suite (cf. Princivalle et al.,
240 2000; Comin-Chiaramonti et al., 2001).

241 The oxygen isotope compositions on separates of clinopyroxene and coexisting olivine ($\delta^{18}\text{O}\text{‰}$)
242 range from 5.5 to 6.0‰, and from 5.0 to 6.1‰, respectively (Table 1 of Comin-Chiaramonti et
243 al., 2001). These measured isotopic ratios are in the range of values for worldwide mantle
244 phases (olivine 4.4 to 7.5‰, clinopyroxene 4.8 to 6.7‰; cf Chazot et al., 1994 and therein
245 references) and for South America mantle xenoliths (olivine 4.9 to 6.4‰, clinopyroxene 5.0 to
246 6.0‰; cf Kyser, 1990). The calculated Cpx-Ol isotopic temperatures (calculated according
247 Kyser et al., 1981) are around 970-1070°C (LK suite) and 1100-1180°C (HK suite), respectively,
248 whereas the Mercier's intracrystalline temperatures are 970-1010°C (LK) and 1005-1050°C
249 (HK), respectively (cf. Petrini et al., 1994; Princivalle et al., 2000).

250 Notably, the $\delta^{18}\text{O}_{\text{Cpx}}$ vs $\delta^{18}\text{O}_{\text{Ol}}$ diagram (Fig. 5A) shows different trends for LK and HK suites,
251 reflecting differences that vary with apparent isotopic equilibration temperatures. In particular
252 the O-isotope temperatures vary coherently with the $\Delta^{18}\text{O}(\text{cpx-ol})$ which tends to 0 in the HK
253 (Fig. 5B). The Mercier's temperatures vs $\Delta^{18}\text{O}(\text{cpx-ol})$ show that HK samples have equilibration
254 temperatures higher than the LK analogues (Fig. 5B). To be noted that the LK samples plot
255 below, but near to the experimentally determined fractionation line (Chiba et al., 1989), whereas
256 the HK-types have less significant O-isotopic fractionation between minerals compared to LK
257 samples (Fig. 5C).

258 According to Princivalle et al. (2000), the investigated Paraguay clinopyroxenes have V(Cell)
259 and V(M1) sites intermediate between those of plagioclase- and garnet-bearing mantle
260 peridotites, i. e, in a pressure range between 12 and 22 kbar (cf. also Nimis and Ulmer, 1998).
261 The ranges of temperature and pressure calculated for the mantle xenoliths from intracratonic
262 Paraguay overlap with those of the mantle xenoliths from the Central Andes, which indicate T
263 900-1190°C and P13-18 kbar (calculated maximum pressures for garnet-free mineral
264 assemblage; Lucassen et al., 2005).

265

266

267 Fig. 5. A

268 **Clinopyroxenes**

269

270 The mantle clinopyroxenes play an important role for the interpretation of mantle processes,
271 such as melting and metasomatism (e.g. Salters and Shimizu, 1988; Takazawa et al., 1992;
272 Vannucci et al., 1994; Laurora et al., 2001. Wang et al., 2007). The decrease of modal
273 clinopyroxene and depletion indexes of whole rock and clinopyroxene (cf Fig. 4), from fertile
274 peridotite (primitive mantle) towards harzburgite-dunite are consistent with the extraction of
275 melt. Melt extraction has a substantial impact on the concentrations of trace elements of the
276 clinopyroxene in general and the incompatible trace elements in particular (cf. Rivalenti et al.,
277 1996). Clinopyroxene is the principal host of important trace elements as Sr, U, Th, Pb and REE
278 in garnet and amphibole free paragenesis (cf. Gregoire et al., 2000).

279 In this section, we examine trace element and Sr and Nd isotopic compositions of the
280 clinopyroxenes from selected Paraguay and Andes mantle xenoliths, LK and HK suites,
281 respectively.

282

283 **Tab. 3.**

284 Fig. 6.

285 Fig. 7.

286 *Trace Elements*

287

288 Trace element analyses are given in Table 3. The main characteristics are summarized in the
 289 extended trace element and REE diagrams of Figs. 6 and 7. The clinopyroxenes from both, LK
 290 and HK suites show variable enrichment of La, Ce, Sr, and Nd, from flat to fairly steep
 291 distribution patterns independently of the Cpx content, (i.e. melting degree of the xenoliths) and
 292 of the calculated temperatures. The orthopyroxene exsolutions (e.g. LK-3254 specimen) and
 293 their host clinopyroxene have similar distribution patterns for the more compatible elements, but
 294 K, LREE and Sr are strongly depleted in opx (Fig. 6C). In particular, the HK-3284
 295 clinopyroxene (Cpx = 1%, with glassy drops up to 7.5%; panels E of Fig. 6 and 7) appears
 296 strongly enriched in REE, comparable to the hosted glassy drops.

297 LREE are commonly strongly enriched (Paraguay: La_N/Yb_N , LK suite, lherzolite = 11.9-5.4,
 298 harzburgite ~ 55; HK suite: lherzolite = 20.3 - 16.2, harzburgite = 98.2 - 7.9), excluding samples
 299 3227 (LK harzburgite with $La_N/Yb_N = 0.92$) and 3222 (HK lherzolite with $La_N/Yb_N = 17.66$
 300 to =0.65; LK Andes $La_N/Yb_N = 110-0.5$; cf. Comin-Chiaramonti et al., 2001, Lucassen et al.,
 301 2005, and Table 3). The contents of incompatible trace elements is not consistent with the
 302 fractional melting processes indicated by the major element composition. Only few samples
 303 have trace element contents consistent with melt extraction (e.g., HK lherzolite 3222 of Comin-
 304 Chiaramonti et al., 2001; panel G of Fig. 6 can be modelled by low degree (6%) of non-modal
 305 fractional melting of primitive mantle; see also Rivalenti et al., 2000).

306 In the normalized trace element diagrams (Fig. 6), all the samples have patterns depleted in Ba,
 307 Nb, Zr and Ti. Furthermore, Zr and Ti display systematically negative spikes (cf. Rampone et
 308 al., 1991), indicating HFSE depletion (cf also Zr/Zr^* and Ti/Ti^* of Table 3). Sr shows both
 309 positive and negative anomalies in LK and HK suites ($Sr/Sr^* = 2.34$ to 0.51), with the

310 minimum in HK harzburgite 3284 ($\text{Sr}/\text{Sr}^* = 0.51$). In particular, Sr from the Andean
311 clinopyroxenes (LK suite) display a positive correlation with $(\text{La}/\text{Yb})_N$, i.e.
312 $\text{Sr}=10.76*(\text{La}/\text{Yb})_N+33,48$ with $r=0.94$ (cf. Lucassen et al., 2005). Y/Y* ratio generally is near
313 the unity (av. 0.94 ± 0.26), but it shows a positive anomaly in the Cpx from LK harzburgite
314 3269 (3.09 to 1.71, core and rim, respectively) and a negative anomaly in Cpx from LK
315 lherzolite 3221 (0.22). In particular, the xenolith A-104 (LK, Andes; Table 3) has a modal
316 composition cpx 17.7, opx 27.3, ol 53.0, sp 2.0%, very similar to that of the primitive mantle
317 (Hofmann,1988; cpx 18, opx 25, ol 55, sp 2%) and the I.E. and REE contents of the
318 clinopyroxene display trends overlapping with those of the PMCE (clinopyroxene trace element
319 compositions; cf. Rivalenti et al., 1996 and panels D of Figs 6 and 7). Moreover, cpx in the
320 residue from non-modal fractional melting (6%) of the primitive mantle mimics the behaviour
321 (minimum contents) of the cpx from HK suite (10-12% cpx; panel G of Fig. 6). It is worth
322 noting that some harzburgites display U-shaped patterns (e.g. LK 3269 and HK 3211; e.g. panel
323 C of Fig. 7), with $\text{La}_N/\text{Sm}_N \sim 60$ and $\text{Gd}_N/\text{Yb}_N \sim 0.55$ (cf Table 3 and Table 2 of Comin-
324 Chiamonti et al., 2001). Such characteristics were attributed to a "carbonatitic component" by
325 Hauri (1997).

326 Compared with the host Cpx, the glassy drops (e.g. HK 3284 Cpx) are strongly enriched in Rb,
327 Ba, and K, and display positive spikes for Sr ($\text{Sr}/\text{Sr}^*=1.679$) and Zr ($\text{Zr}/\text{Zr}^*=1.13$; cf. panel E of
328 Fig. 6, and Table 3). Notably, the geochemical features of the glassy drops fit those of a melt
329 from a phlogopite rich source (e.g. McKenzie and O'Nions, 1991: Ti = 4796 ppm and K= 78786
330 ppm). The glassy patches differ from the glassy drops by their Zr/Zr^* (0.47 vs 1.13), Ti/Ti^*
331 (0.16 vs 0.69), Eu/Eu^* (1.14 vs 0.83), and by a positive Nb spike (panel F of Fig. 6). The
332 general element distribution pattern of the glassy patches is similar to that of glasses, which
333 were interpreted by Laurora et al. (2001) as the disequilibrium melting of metasomatic
334 assemblages (dominantly amphibole) that produced carbonated silicate melts. Correlations
335 between $\delta^{18}\text{O}$ and IE contents were not observed (cf. Comin-Chiamonti et al., 2001): the lack

336 of significant differences in oxygen isotopic composition between LREE-depleted and REE-
337 enriched samples of both LK and HK suites suggests that metasomatism does not affected the
338 oxygen isotope systematics.

339 Clinopyroxenes from Andean pyroxenite are characterized by element contents of about 1-2
340 times primitive mantle (Fig. 6, panel H), but with a prominent negative spike for Sr (Sr/Sr^*
341 0.42-0.45) and Y (Y/Y^* 0.67-0.78), and a slight upward-concave REE pattern (Fig. 7, panel H
342 and Table 3). Clinopyroxene from carbonatized xenoliths display a large range of the
343 normalized element contents between 0.1 to 10 times primitive mantle (Fig. 6, panel H) and
344 variable REE patterns (Fig. 7, panel H).

345 In summary, the trace elements patterns of clinopyroxenes from Paraguayan and Andean
346 xenoliths could be the result of different, superimposed processes affecting the upper mantle (cf:
347 Comin-Chiaramonti et al., 2001):

348 depletion by variable degrees of partial melting, as shown by the "residual" character of
349 mantle xenoliths;

350 subsolidus re-equilibration, as indicated by spinel and orthopyroxene exsolution in
351 clinopyroxenes, and subsequent variations of the $V(\text{cell})$ and $V(\text{M1})$ volumes (Pricivalle
352 et al., 2000);

353 metasomatic processes, as shown by variable enrichment of incompatible elements, both
354 in the clinopyroxenes and in the host xenoliths;

355 decompression and partial melting after the sampling by host lavas, as shown by the
356 presence of glassy phases, i.e. glassy drops in clinopyroxenes "cutting" the spinel
357 exsolutions, and by the blebs in the host xenoliths.

358

359 **Tab. 4.**

360

361 *Sr-Nd Isotopes*

362

363 Rb-Sr and Sm-Nd isotope data have been determined on whole rock and on clinopyroxene-
364 orthopyroxene separates (Table 4). Whole rocks and mineral separates from Paraguay display a
365 very wide isotopic variation (measured ratios), with $^{87}\text{Sr}/^{86}\text{Sr}$ ratios overlapping in the LH and
366 HK suites and in the ranges 0.70360 - 0.70480 (WR) and 0.70326 - 0.70461 (Cpx), respectively,
367 and $^{143}\text{Nd}/^{144}\text{Nd}$ in the ranges 0.51246 - 0.51346 (WR) and 0.51260 - 0.51348 (Cpx),
368 respectively. In the Andean Cpx are mainly in the ranges 0.70271-0.70478 ($^{87}\text{Sr}/^{86}\text{Sr}$) and
369 0.51280-0.51387 ($^{143}\text{Nd}/^{144}\text{Nd}$) and compositions from both distant areas overlap.

370 In Paraguay, positive correlation ($r^2 > 0.8$) between $^{87}\text{Rb}/^{86}\text{Sr}$ and $^{87}\text{Sr}/^{86}\text{Sr}$ is shown by
371 xenoliths (whole-rocks) of the two suites, from which a reference lines of 128 Ma can be
372 calculated for the two suites (cf. inset Fig. 6B of Comin-Chiaramonti et al., 2001). $^{87}\text{Sr}/^{86}\text{Sr}$ and
373 $^{87}\text{Rb}/^{86}\text{Sr}$ ratios of the pyroxenes are poorly correlated with the 128 Ma reference line, reflecting
374 substantial isotopic disequilibria between whole rock and hosted clinopyroxene (cf. Song and
375 Frey, 1989). The $^{87}\text{Sr}/^{86}\text{Sr}$ xenolith whole rock - clinopyroxene disequilibria could be ascribed
376 to contributions of a component enriched in radiogenic Sr to the bulk xenoliths (cf. Downes and
377 Dupuy, 1987). This component actually may be confined to the blebs, to glassy drops, and/or
378 glass at grain boundaries. The leaching of the whole rock samples possibly was not enough
379 strong to dissolve such an enriched phase, as in the case of glassy component of the
380 clinopyroxene. The leaching solutions produce errorchron lines (Table 4 of Comin-
381 Chiaramonti et al., 2001) between 129 Ma (LK sample 3192) and 61 Ma (HK sample 3307),
382 approaching the ages of the K-alkaline magmatism (128 Ma) and nephelinitic magmatism of
383 the region (60 Ma), respectively (Velázquez et al., 2006). Probably, an isotopic mixing between
384 an old whole-rock component (?) and the magmatic-decompressional event (Early Cretaceous to
385 Paleogene) affected the clinopyroxenes and the host xenoliths.

386 The Early Cretaceous age approaches those inferred for potassic and tholeiitic rocks (128 and
387 133 Ma, respectively; Comin-Chiaramonti and Gomes, 1996, 2005; Ernesto et al., 1996) from
388 Asunción-Sapucaí-Villarrica graben. Notably, the Early Cretaceous age may indicate the latest
389 thermal event affecting the lithospheric mantle before the sampling by ascending
390 melanephelinite magma in the late Early Cretaceous to Eocene times in Eastern Paraguay. In the
391 Andes the age of equilibration of the lower crust xenoliths and the eruption of the host lavas is
392 around 100 Ma (Lucassen et al., 2005). Thus measured isotope ratios of the analyzed samples
393 (Table 4) were corrected for in situ decay to ages of 128 and 100 Ma for Paraguay and Andes,
394 respectively.

395 The initial $^{87}\text{Sr}/^{86}\text{Sr}$ and $^{143}\text{Nd}/^{144}\text{Nd}$ ratios of whole rocks and clinopyroxenes (Fig. 8)
396 encompass the whole range of isotopic ratios observed in NE Brazil (Pico Cabuji), Fernando de
397 Noronha island and Patagonia xenoliths (cpx compositions; Rivalenti et al., 2000; Barbieri et al.,
398 1997; Stern et al., 1989). The xenoliths compositions partly overlap with the ratios of their host
399 rocks (Fig. 8). $^{87}\text{Rb}/^{86}\text{Sr}$ ratios of the clinopyroxenes are systematically lower than those of their
400 host rocks, and cluster between 0.0002 and 0.0086; only one sample from LK suite and two
401 samples from HK suite have $^{87}\text{Rb}/^{86}\text{Sr}$ ratios above 0.1 (cf Table 3 of Comin-Chiaramonti et al.,
402 2001 and Table 3 of Lucassen et al., 2005).

403 The initial $^{143}\text{Nd}/^{144}\text{Nd}$ ratios span from enriched (0.51237) to depleted (0.51372) signatures
404 (Table 4; Fig.8). The Nd isotope ratios of the Paraguay - Andes xenoliths and hosted pyroxenes
405 encompass almost the entire range of values reported for mid ocean ridge basalts (MORB) and
406 ocean island basalts (OIB) (cf White, 1985; Galer and O'Nions 1989; Stosch and Lugmair,
407 1986). Significant correlations between $^{147}\text{Sm}/^{144}\text{Nd}$ and $^{143}\text{Nd}/^{144}\text{Nd}$ were not observed ($r^2 \leq$
408 0.85), but the best fitting is along a 0.5 Ga line (cf. Table 4 of Comin-Chiaramonti et al., 2001).
409 On the contrary, strongly positive correlations are shown by the clinopyroxenes from Pico
410 Cabuji (NE Brazil; $r^2 = 0.99$, best fitting at 1110 Ma) and from Fernando de Noronha ($r^2 = 0.90$,

411 best fitting at 413 Ma; cf Table 4 and Fig. 6B of Comin-Chiaramonti et al., 2001). At any way,
 412 to be noted that the Paraguay xenoliths tend to fit the “Paraguay array” of Comin-Chiaramonti
 413 et al., (1997). In many of the Andean $^{147}\text{Sm}/^{144}\text{Nd}$ (too low or 'enriched') do not fit the observed
 414 $^{143}\text{Nd}/^{144}\text{Nd}$ (too radiogenic) ratios and young enrichment is needed in the Andean case to
 415 explain these relationships. The poor fit of $^{143}\text{Nd}/^{144}\text{Nd}$ and $^{147}\text{Sm}/^{144}\text{Nd}$ for clinopyroxenes
 416 could have variable causes, initial Nd isotope heterogeneity, variable ages or intensity of the
 417 metasomatic overprint or depletion by melt extraction.

418

419 **Fig. 8**

420

421 In particular, the $^{143}\text{Nd}/^{144}\text{Nd}$ ratios of clinopyroxenes and orthopyroxene indicate isotopically
 422 heterogeneous mantle sources, ranging from values approaching the Bulk Earth estimate (e.g.
 423 HK-3288-Cpx: measured $^{143}\text{Nd}/^{144}\text{Nd} = 0.51267$) to a variably LREE depleted source (e.g. LK-
 424 3254-Cpx: $^{143}\text{Nd}/^{144}\text{Nd} = 0.51348$). Sample 3252 (LK) shows significant whole rock-
 425 clinopyroxene-orthopyroxene isotopic disequilibria, i.e. measured $^{143}\text{Nd}/^{144}\text{Nd} = 0.51311$,
 426 respectively, and $^{147}\text{Sm}/^{144}\text{Nd} = 0.16$, 0.11 and 0.10, respectively (cf: Table 4). Thus, the
 427 decoupling between $^{143}\text{Nd}/^{144}\text{Nd}$ and $^{147}\text{Sm}/^{144}\text{Nd}$ ratios suggests a complex history of REE
 428 depletion and enrichment events in the mantle, as envisaged by the trace element patterns.

429 The higher $^{87}\text{Sr}/^{86}\text{Sr}$ ratios for whole rocks, and acid leachates relative to clinopyroxenes
 430 indicate that the Sr isotopic ratio was affected by secondary processes. Therefore, model age
 431 (depleted mantle) calculations based on Rb/Sr systematics are not reliable. Sm/Nd model ages
 432 (TDM), calculated with respect to a depleted reservoir (Table 4), range from 109 to 921 Ma
 433 (average LK Paraguay = 0.46 ± 0.26 Ga, HK, 0.50 ± 0.12 Ga, and LK-Andes 0.49 ± 0.09 Ga),
 434 except the pyroxenite for which a model age of ~ 1410 Ma was calculated. Unfortunately, Nd
 435 model ages are highly problematic in metasomatically overprinted rocks, which are seen in the
 436 mismatch of $^{147}\text{Sm}/^{144}\text{Nd}$ and $^{143}\text{Nd}/^{144}\text{Nd}$ ratios. The approach relies on a refractory Sm/Nd

437 system after separation of the respective batches from the depleted mantle or requires a good
438 control (time and composition) of changes in the Sm/Nd system in the respective batch.

439 The Sm/Nd model ages (TDM), calculated with respect to a depleted reservoir (Table 4), range
440 from 110 to 920 Ma (average LK = 0.46 ± 0.26 , HK, 0.50 ± 0.12 Ga, respectively, where LK-
441 Andes show 0.49 ± 0.09 Ga, clustering the Brasiliano cycle, i.e. 900-460 Ma, according to
442 Hartmann et al., 2000), pyroxenite excepted for which a model age of 1413 Ma was calculated.
443 The complexities of the Sm/Nd systematic suggests a multi-stage history, such as melt
444 extraction and fluid or melt addition, which all could be highly variable and occur repeatedly
445 and were preserved in the lithospheric mantle beneath Paraguay and Andes at different times.

446

447 *Lead Isotopes*

448

449 Lead isotopes are available for the clinopyroxenes from Andes (Lucassen et al., 2005) and for
450 two whole rock samples from Paraguay (Table 5). Xenoliths from both distant areas, Paraguay
451 and the Central Andes, show similar Pb isotope compositions (Fig. 9). The $^{206}\text{Pb}/^{204}\text{Pb}$,
452 $^{207}\text{Pb}/^{204}\text{Pb}$, $^{208}\text{Pb}/^{204}\text{Pb}$ ratios vary between 17.7 and 20.6, 15.5 and 15.7, 38.1 and 41.9,
453 respectively, but the majority of the samples cluster in the much smaller field of their host lavas
454 and the Bolivian Ayopaya carbonatite (Schultz et al., 2004). The $^{238}\text{U}/^{204}\text{Pb}$ (μ ratio) is below
455 the typical depleted mantle value of ca. 8 (except sample A-147-m with $\mu \sim 55$). There is again
456 as in the case of the Sm-Nd isotope system a mismatch between parent/daughter ratios and the
457 radiogenic isotope composition (Table 5) and a metasomatic overprint or melt extraction
458 relatively late in the history of the lithospheric mantle is required.

459 The isotopic compositions of many samples, roughly plot in between EMI, DMM and HIMU
460 mantle end members (Fig. 9; cf. Zindler and Hart, 1986; Stracke et al., 2003) and could be
461 tentatively explained by a mixing scenario (metasomatism), which adds a HIMU signature from

462 an external source (convective mantle) to a EM1 lithosphere (cf. Figs. 3 and 9). However, there
463 are also a number of samples, which do not fit such a simple trend.

464

465

466

Table 5.

467

Discussion and Conclusive remarks

468

469 Paraguay and Andes spinel peridotites, which occur at similar latitudes, are variable in major
470 element compositions, ranging from relatively "fertile" to very depleted in basaltic component.

471 Two suite can be distinguished: low and high in K_2O content ($K_2O < 0.1$ and >0.1 , LK and HK,
472 respectively). The HK suite is restricted to xenoliths from Paraguay and some of these xenoliths

473 have exceedingly high K_2O and incompatible elements (IE) compared to the composition of

474 lherzolites which underwent partial melting during basalt-extraction (cf. Frey and Green, 1974).

475 The major element of the xenoliths are largely consistent with residual compositions after
476 variable degrees of partial melting, but metasomatic effects are apparent above all in the HK

477 suite. The incompatible element contents of clinopyroxenes from LK and HK is highly variable

478 similar to other world-wide occurrences (cf. e.g., Salters and Shimizu, 1988). Trace element

479 enrichment suggests that some processes, other than depletion, played an important role in the

480 xenoliths (cf. Fig. 6).

481 Paraguayan and Andean clinopyroxenes display very variable REE enrichments (Fig.7), which

482 are not bound to a specific major element composition or mineralogy. Only in few Paraguayan

483 samples, the enrichments appear to be more pronounced in clinopyroxene with spongy textures

484 and glassy drops. Considering the radiogenic Nd isotope compositions and the low Sm/Nd

485 ratios in many clinopyroxene, this enrichment occurred probably late in the history of the

486 xenoliths and it affected xenoliths with different isotopic signatures, i.e. equally with radiogenic

487 and relatively unradiogenic isotope signatures. The U-Pb system indicates also a late

488 metasomatic event, with the removal of U from the rocks (or enrichment of Pb). The radiogenic
489 isotope signatures of the xenoliths do not fit any of the canonical mantle end-members. They
490 span a broad compositional field, most samples with Sr and Nd isotope compositions indicating
491 the contribution of depleted mantle, but some samples show also affinity to enriched mantle of
492 the EMI-type typical of old subcontinental mantle known from the Brazilian Shield (cf. Fig. 8).
493 Pb isotope compositions are also variable, but most samples are more radiogenic than average
494 depleted mantle and indicate the influence of an HIMU source (Fig. 9). The magmatic host
495 rocks of the xenoliths have a broadly similar isotopic composition (cf. Figs 3 and 9), but some
496 xenoliths samples show extreme compositions.

497 There are two hypothetical end-member scenarios for creating heterogeneous isotope
498 compositions in the lithospheric mantle sources: (1) radiogenic in-situ growth during long-term
499 separation from convection and (2) metasomatism of the lithospheric mantle by infiltrating
500 melts from an external source, i.e. the asthenosphere (depleted mantle) or deeper mantle sources
501 (e.g. HIMU). The processes must be pervasive (the signatures are found in xenoliths from the
502 uppermost mantle and in the host melts from below) and at large scale, because at least in the
503 Andes a large N-S section shows similar signatures.

504 Preconditions of (1) are the existence of a reasonably old lithosphere, which formed in the
505 Palaeozoic or Pre-Palaeozoic and ancient substantial reworking and metasomatism of this
506 lithosphere. The pristine material of mantle lithosphere is assumed to be of asthenospheric
507 (depleted mantle) origin. Substantial reworking probably took place at Early Palaeozoic for the
508 Andean and similar or older for the Paraguayan section. Such scenario has been invoked for
509 compositional unusual, high- μ intra-plate magmatism on the Arabian plate (Stein et al., 1997).

510 The second scenario includes infiltration of the pristine lithosphere by melts from an external
511 high- μ source. This event(s) must be relatively young, due to the already highly radiogenic Pb
512 compositions of the required high- μ agent. Further, this scenario requires large scale material

513 transport from the deeper mantle through the asthenosphere and hence a young convective
514 (thermal) instability for the activation of the high- μ reservoir.

515 However, the second model creates problems at least with the Andean intra-plate magmatic
516 rocks and xenoliths. Contemporaneous intense magmatism in the magmatic arc is clearly
517 derived from a subduction modified depleted mantle, i.e. involves 'normal' asthenosphere
518 without any contribution of a HIMU component (e.g. Lucassen et al., 2006). The assumption of
519 a plume-like scenario for the intra-plate magmatism and metasomatism in the lithospheric
520 mantle would require two separated, compositionally different large scale convection systems,
521 which did not mix on a restricted space (distance Mesozoic arc - intra-plate magmatism is 500
522 present + 200 km Andean shortening \sim 700 km in the Cretaceous). Assuming an older 'plume'
523 event this must have been without important magmatic activity. There are no indicators of a
524 large scale melt infiltration/melting event involving HIMU in the magmatic record. The last
525 thermal pulse recorded on a large N-S scale also far behind the arc was late Carboniferous-
526 Permian. Mesozoic to Cenozoic magmatism was very minor and the Salta Rift (and other
527 extensional structure) was considered as of minor magmatic activity.

528 The first, lithospheric scenario has the obvious advantage that it does not require additional far
529 reaching geodynamical implications. In a first step in-situ growth of the variable isotope
530 signatures is caused by variable old metasomatism (trace element signatures), in the case of the
531 Andes linked to the Early Palaeozoic orogeny. Incompatible trace element distribution is easily
532 changed by minor amounts of melting, and trace elements could be locally redistributed by trace
533 element rich melts, with or without changes of the isotope composition. Therefore, the late
534 metasomatism causing the mismatch of parent element contents and daughter isotope
535 compositions could be largely autochthonous and caused by Na-alkaline basaltic magmas from
536 deeper, possibly garnet-bearing lithospheric mantle (cf. Comin-Chiaramonti et al, 1997).
537 Moreover, high LREE and Sr abundances coupled with depletion in Nb, Ti, Zr of some
538 xenoliths from Paraguay could indicate carbonatite metasomatism (Comin-Chiaramonti et al.,

539 1991; Hauri et al. 1993). In Paraguay the situation is complicated by the occurrence of old
540 subcontinental mantle of the Brazilian Shield, which could contribute to possible redistribution
541 of isotope and trace element signatures by melts within the lithospheric mantle. Unfortunately
542 little of such processes -except compositional heterogeneities- remained visible in high T
543 environments, due to rapid recrystallization. The melt inclusions of some Paraguayan xenoliths
544 are the only preserved visible heterogeneities linked to metasomatism.

545 In Paraguay, the isotopic data seem to indicate that the lithospheric mantle prior to the
546 enrichment event was dominated by a depleted component, isotopically resembling MORB
547 sources or even more depleted represented by residua from ancient events of partial melting. In
548 the Andes, the composition of the source is also believed to be depleted mantle, but different
549 from a MORB type source, showing slightly more radiogenic ^{87}Sr and less radiogenic ^{143}Nd
550 than the Pacific MORB. The substantial compositional variations in some xenoliths and host
551 lavas likely reflect small-scale sampling of the lithospheric mantle, and differential interactions
552 between fluids and overlying peridotites.

553 In summary, there is strong evidence for variable metasomatism in the xenoliths in both distant
554 areas. The observed radiogenic isotope trend (Bulk Earth vs Depleted Mantle) is not consistent
555 with major element refractory parameters, suggesting that a mixing with enriched components is
556 also recorded at whole-rock scale.

557 We stress that the compositions of mantle xenoliths and host sodic alkaline lavas are
558 surprisingly similar at about 26°S latitude in an extensional intra-cratonic environment
559 (Paraguay) and an extensional to transpressional setting transitional between arc and craton
560 (Andes orogen).

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1

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6

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208

1

CAPTIONS TO THE FIGURES

2 **Fig. 1. A:** Map of the studied regions (modified after Comin-Chiaramonti et al., 2007, Lucassen et al., 2007) showing
 3 contours (red lines) of the depth (km) of the subducting Nazca slab based on seismic data (Gudmundsson and
 4 Sambrige, 1998). Heavy lines (black) outline the Cretaceous rift systems. The hatched area roughly marks the
 5 extension of intense early Paleozoic reworking of Proterozoic material, but the exact border to the Brazilian Craton
 6 remains unknown (cf. Lucassen et al., 2000). Pink fields delineate inferred positions of major cratonic fragments
 7 below Phanerozoic cover (after Laux et al., 2005): AAB, Arequipa-Antofalla; AC, Amazon Craton; AB, Apa
 8 Block; PR, Paranapanema; LP, Rio de la Plata; PA, Pampia. Localities characterized by sodic alkaline magmatism,
 9 mantle xenoliths bearing: 1, Asunción; 2, Misiones; 3, Belén; 4, Las Conchas and Cadillal; 5, Finca del Rodeo; 6,
 10 Betanzos (data sources: Comin-Chiaramonti et al., 2007 and Lucassen et al., 2007).

11 **B:** Seismic tomography image of Liu et al. (2003) across a profile approximately at 24° Lat. S. Note that the low-
 12 velocity feature in the mantle to the East has been interpreted as a fossil mantle plume by VanDecar et al. (1995).

13

14 **Fig. 2.** Plot, following De la Roche's (1980, 1986) nomenclature, of the rock-types reported in Table 1. $R_1=4Si-11(Na+K)-2(Fe+Ti)$, $R_2=6Ca+2Mg+Al$; 1, ankaratrite; 2, basanite; 3, alkali basalt; 4, nephelinite; 5, tephrite; 6,
 15 trachybasalt. The whole compositional fields for the Misiones (M, light grey) and Asunción (A, dark grey) are
 16 also shown. **Inset:** Na_2O vs K_2O (wt%) diagram. Data source: Velázquez et al. (2006) and therein references;
 17 Lucassen et al. (2007).

18

19
 20 **Fig. 3. A:** Trace elements are normalized to the primitive mantle (Sun and McDonough, 1989) for average
 21 compositions of the mafic alkaline rocks from the Paraguay and Andes compared with the whole field of the
 22 Paraguay mafic (sodic) analogues. **B:** REE are normalized to chondrite (Boynnton, 1984) Data source as in Fig. 2.

23 **C, D, E:** Isotopic mixing curves between HIMU and potassic magmas from Early Cretaceous potassic rocks
 24 (ECK, Eastern Paraguay), computed using the following isotopic composition: HIMU (St. Helena; Chaffey et al.,
 25 1989); $^{87}Sr/^{86}Sr=0.70282$, $Sr=650$, $^{143}Nd/^{144}Nd=0.5130$, $Nd=40$, $^{206}Pb/^{204}Pb=20.73$, $^{207}Pb/^{204}Pb=15.77$ and
 26 $^{208}Pb/^{204}Pb=40.80$, $U=1.44$, $Th=3.88$, $Pb=4$, $\mu=24.4$, $\kappa=2.78$; ECK, A: $^{87}Sr/^{86}Sr=0.7070$, $Sr=1300$,
 27 $^{143}Nd/^{144}Nd=0.5117$, $Nd=60$, $^{206}Pb/^{204}Pb=16.672$, $^{207}Pb/^{204}Pb=15.422$, $^{208}Pb/^{204}Pb=37.10$, $U=1.47$, $Th=6.38$,
 28 $Pb=2$, $\mu=23.09$, $\kappa=4.80$; ECK B: $^{87}Sr/^{86}Sr=0.7070$, $Sr=1300$, $^{143}Nd/^{144}Nd=0.5117$, $Nd=60$, $^{206}Pb/^{204}Pb=16.945$,
 29 $^{207}Pb/^{204}Pb=15.434$, $^{208}Pb/^{204}Pb=37.369$, $U=2.40$, $Th=9.10$, $Pb=15$, $\mu=9.8$, $\kappa=3.8$. Crosses: 10% step of mixing.
 30 DMM, HIMU and EMI components after Hart and Zindler (1989). Crystalline basement as in Antonini et al.,
 31 2005. For comparison the Ayopaya carbonatite is plotted (Lucassen et al., 2007).

32

33 **Fig. 4. A.** $(SiO_2 + Al_2O_3)/(FeO_t + MgO)$ vs $CaO/(FeO_t + MgO)$ diagram (molar ratios) for bulk-rock compositions
 34 relative to the Paraguay xenoliths, LK (I), HK (II) and Andean (III) suites, respectively. Stars B, H and BE:
 35 pyrolyte compositions (Bristow, 1984), primitive mantle (Hofmann, 1988) and Bulk Earth (McKenzie & O'Nions,
 36 1991), respectively. Data sources: Demarchi et al. (1988) and Lucassen et al. (2005). **B:** variation of the modal
 37 cpx vs the modal cpx/opx ratio. The line indicates model variation trends induced by non-modal fractionation
 38 melting in a primitive mantle composition (cf. Rivalenti et al., 2000); thicks are at 4% melting intervals; MM,
 39 mantle mode after Rivalenti et al., 1996.

40

41 **Fig. 5. A -** $\delta^{18}O_{Cpx}$ vs $\delta^{18}O_{Ol}$; regression lines for LK and HK suites and equilibration line ($\Delta = 0$), respectively, are
 42 shown. **B -** $\Delta^{18}O_{Cpx-Ol}$ vs isotopic equilibration temperatures, T°C, calculated according to Kyser (1990); the
 43 grey field represents the clinopyroxene-olivine pairs from South America mantle xenoliths (Kyser, 1990). **C -**
 44 Clinopyroxene-olivine fractionation as a function of clinopyroxene intracrystalline temperatures (Mercier, 1980);
 45 Chiba: fractionation line from Chiba et al. (1989). The data relative to the O isotopic values (V-SMOW9 are also
 46 shown; data source Comin-Chiaramonti et al., 2001.

47

48 **Fig. 6.** Trace element patterns of clinopyroxenes from Nemy xenoliths, normalized to the primitive mantle (Hofmann,
 49 1988) subdivided following LK (Py, Paraguay; An, Andes) and HK (Py) typology (panels A to D: LK; panels E
 50 to G: HK) and content % of the clinopyroxenes. Cpx, clinopyroxene, Opx, orthopyroxene, GD: The patterns of
 51 glassy drops and glassy patches are also shown. PMCE: clinopyroxene trace element composition in primitive
 52 mantle (Hofmann, 1988; cf. Rivalenti et al., 1996). In panels D and F the clinopyroxene pattern in the residue
 53 from non-modal fractional melting (6%) of primitive mantle (Hofmann, 1988) is reported; fractional melting
 54 parameters and partitions coefficients, after Johnson et al. (1990) and Rivalenti et al. (1996), respectively. Data
 55 sources, as in Table 3.

56 **Fig. 7.** REE patterns of clinopyroxenes normalized to the primitive mantle (Hofmann, 1988), and REE normalized to
57 chondrites (Boynton, 1984). Symbols and panels, as in Fig. 6.

58

59 **Fig. 8.** $^{87}\text{Sr}/^{86}\text{Sr}$ vs $^{143}\text{Nd}/^{144}\text{Nd}$ plot for Ñemby clinopyroxenes and host mantle xenoliths; for comparison the
60 compositions of the cpx from NE Brazil (NEB) and Fernando de Noronha (FDN; Rivalenti et al., 2000),
61 Patagonia (PAT; Stern et al. 1989, Barbieri et al., 1997) are also plotted. L, leaching solutions (Comin-
62 Chiamonti et al., 2001). A, Host lavas from Andes (Lucassen et al., 2007); P, host lavas from Paraguay
63 (Comin-Chiamonti and Gomes, 1996, 2005).

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65 **Fig. 9.** Sr, Nd and Pb isotopic composition of clinopyroxene separates and two Paraguay mantle xenoliths (cf. Fig. 3).
66 Grey field: host lavas. Data sources as in Table 5. Symbols as in Fig. 3.

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CAPTIONS TO THE TABLES

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Table 1. Averaged compositions (in brackets standard deviations) for the sodic alkaline rock-types from Eastern Paraguay and Central Andes. Numbers as in Fig. 1. Data sources: Comin-Chiaramonti et al. (2001), Velázquez et al. (2006), Lucassen et al. (2007). (Eu/Eu)* calculated according to the chondritic values after Boynton, 1984.

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Table 2. Averaged isotopic analyses, time integrated ratios. For comparison the Apoyaya data are shown (Schultz et al., 2004 and Lucassen et al., 2007). Other sources as in Table 1.

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Table 3. Trace elements in pyroxenes (Max and min, maximum and minimum concentration) from Paraguay (A) and Andes (A) mantle xenoliths, LK and HK suites, respectively and for glassy drops in clinopyroxene and glassy patches in xenoliths. Data sources: Petrini et al. (1994), Comin-Chiaramonti et al. (1992, 2001), Lucassen et al. (2005). Primordial mantle Normalized ratios, PM; chondrite normalized ratios, CH. Paraguay data, ion microprobe analyses (analytical methods in Petrini et al., 1994), Andes data, ICP-MS (analytical methods in Lucassen et al., 2007).

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Table 4. Rb, Sr, Sm, Nd contents and measured, initial ratios for Sr and Nd isotope ratios, ordered for increasing content of clinopyroxene (Cpx%). Data sources: Petrini et al. (1994), Comin-Chiaramonti et al. (2001), Velázquez et al. (2006), Comin-Chiaramonti and Gomes (2005), Lucassen et al. (2005). For comparison, data relative to the Apoyaya carbonatite (as in Lucassen et al. 2007) are reported. The following notional ages were considered for in-situ decay (s. text): Paraguay, 128 Ma, Andes, 100 Ma, Apoyaya, 140 Ma. LK, HK, low- and high K suites, respectively; MIS, misiones; ASU, Asunción; WR, whole rock, CPX, clinopyroxene, OPX, orthopyroxene; CARB, carbonatized peridotite xenoliths, PYROX, pyroxenitic xenoliths. TDM: calculation of model dates relative to a depleted reservoir, $^{143}\text{Nd}/^{144}\text{Nd} = 0.513114$ and $^{147}\text{Sm}/^{144}\text{Nd} = 0.222$; Faure, 1986) are presented for plausible solutions.

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Table 5. U, Th, Pb abundances and Pb isotopic composition of Andean clinopyroxenes and Paraguay xenoliths. Data sources: Lucassen et al. (2005), Comin-Chiaramonti and Gomes (2005). Only specimens with available isotopic Sr-Nd initial ratios are reported (cf. Table 4).

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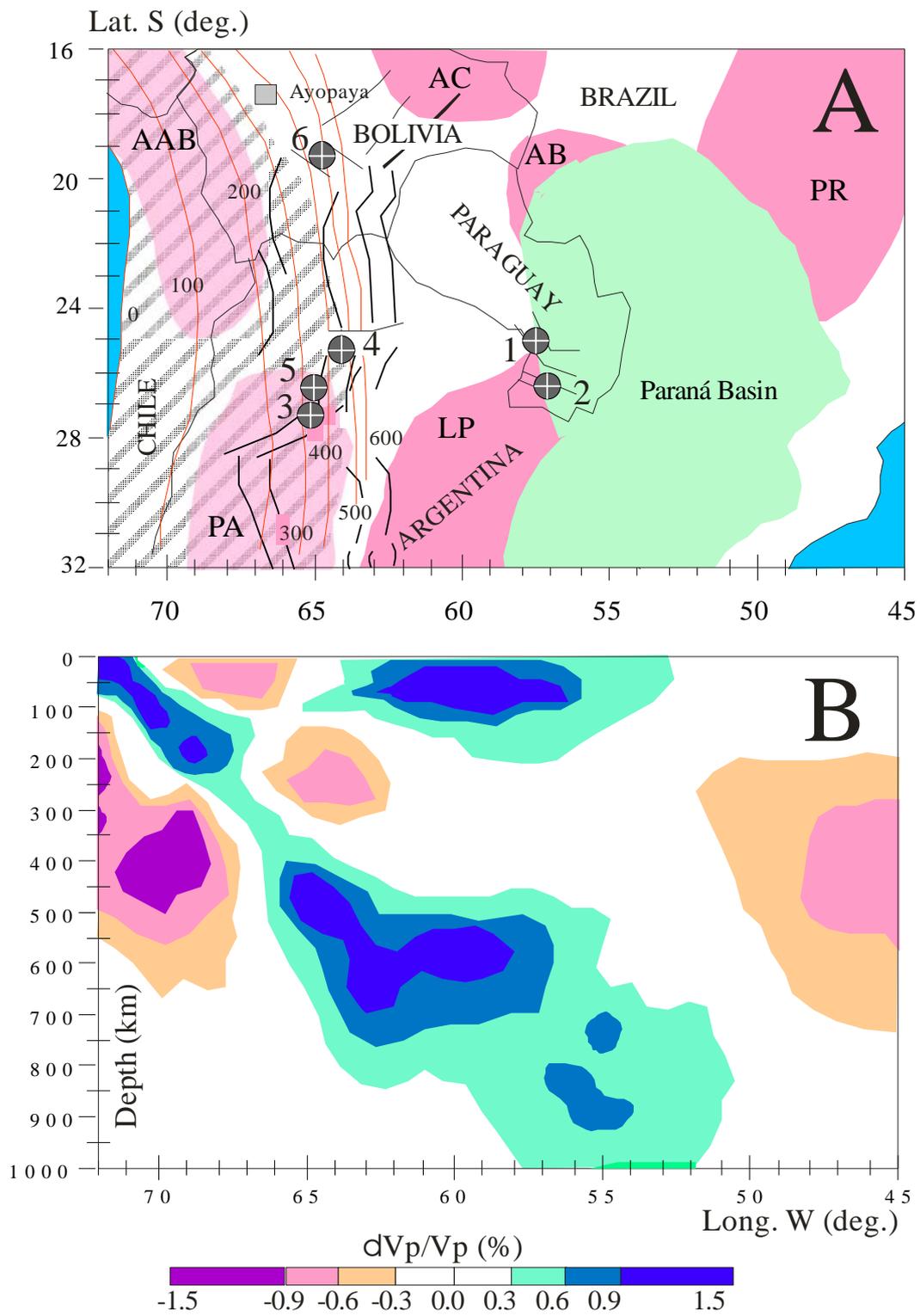


Fig.1

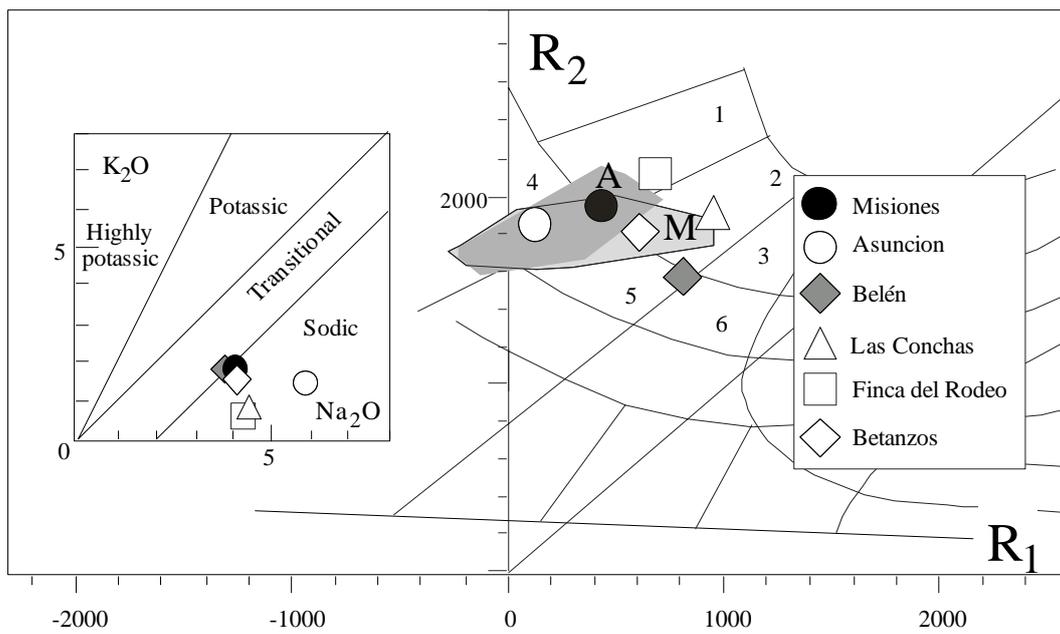


Fig. 2

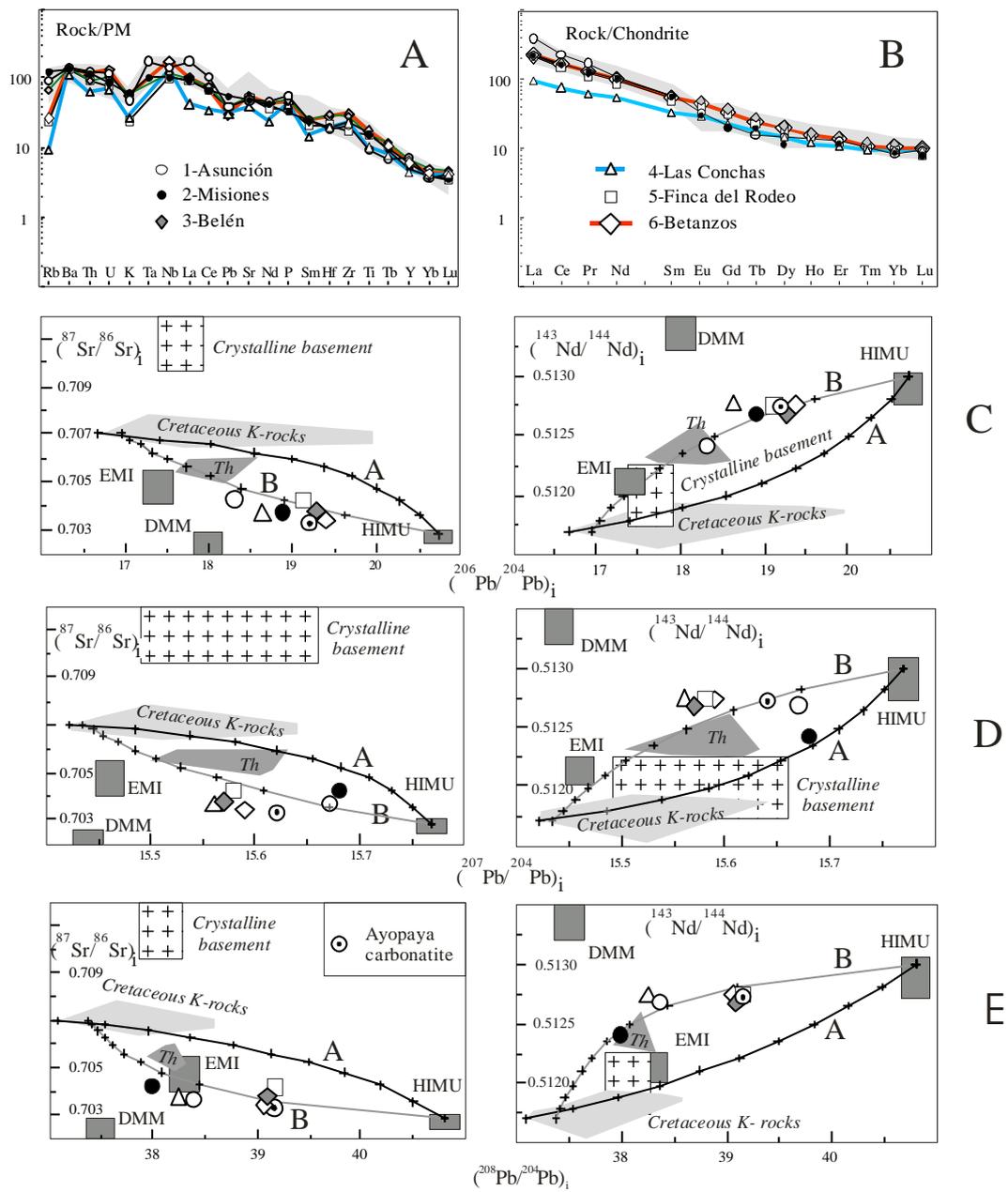


Fig. 3

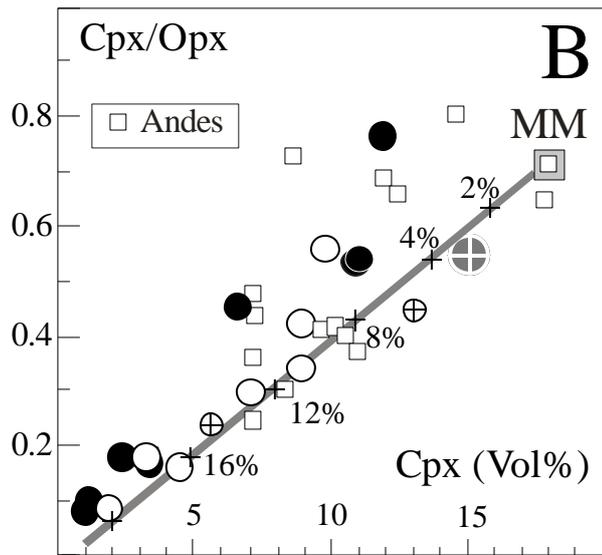
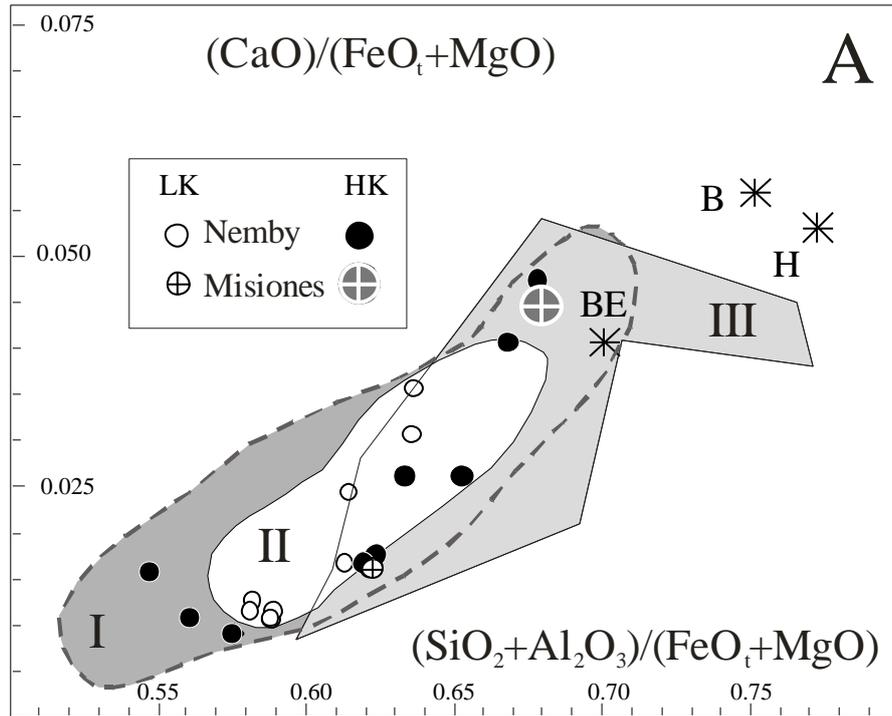


Fig. 4

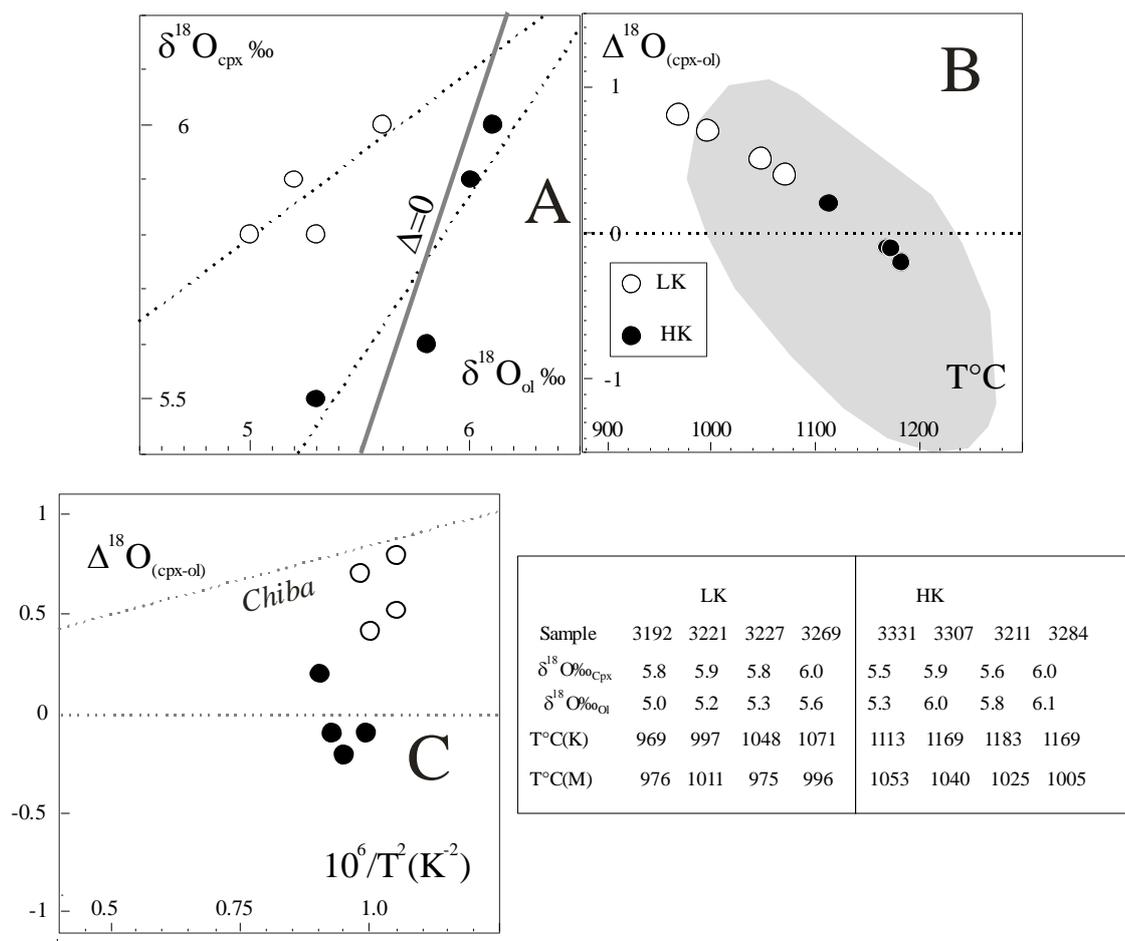


Fig. 5

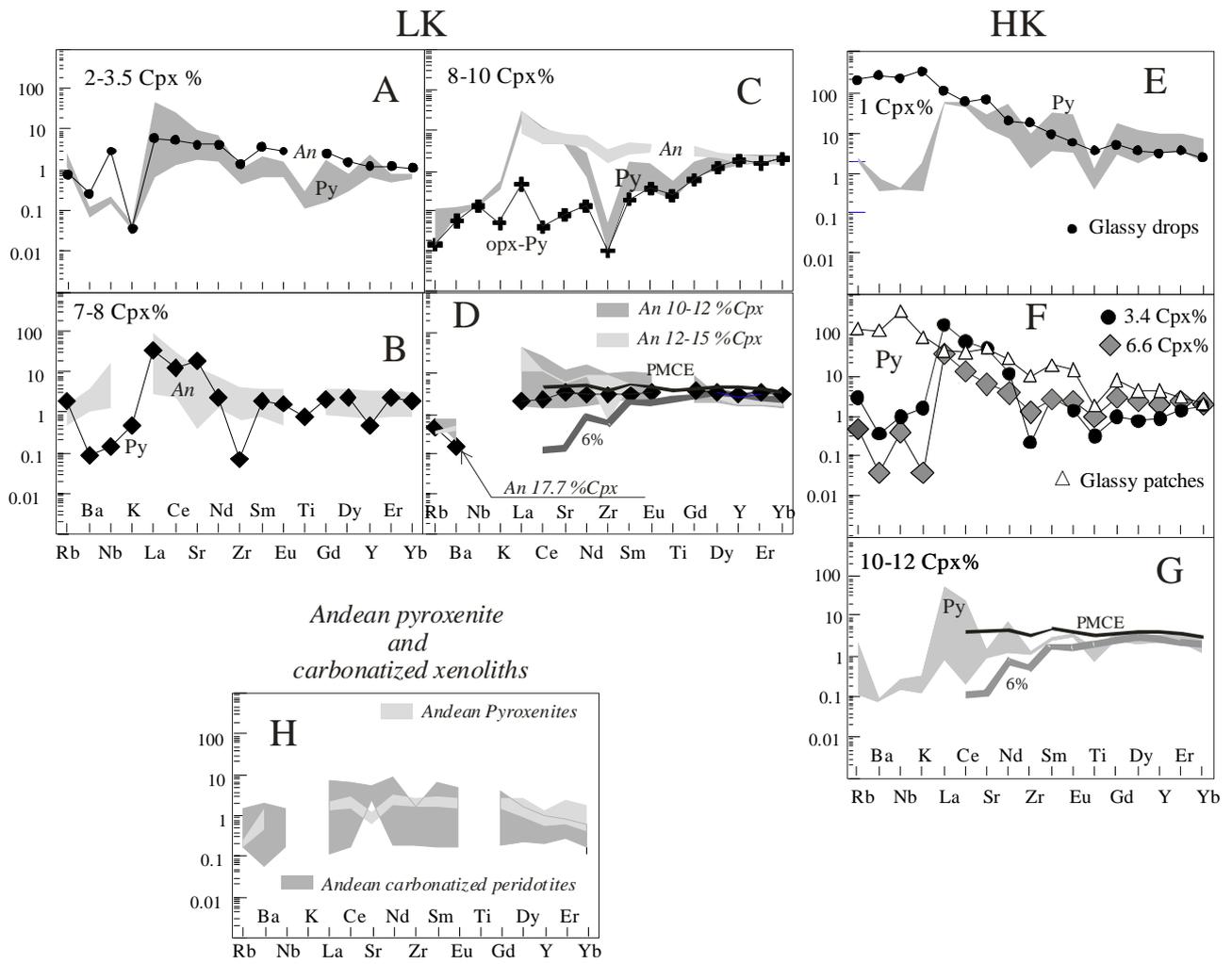


Fig. 6

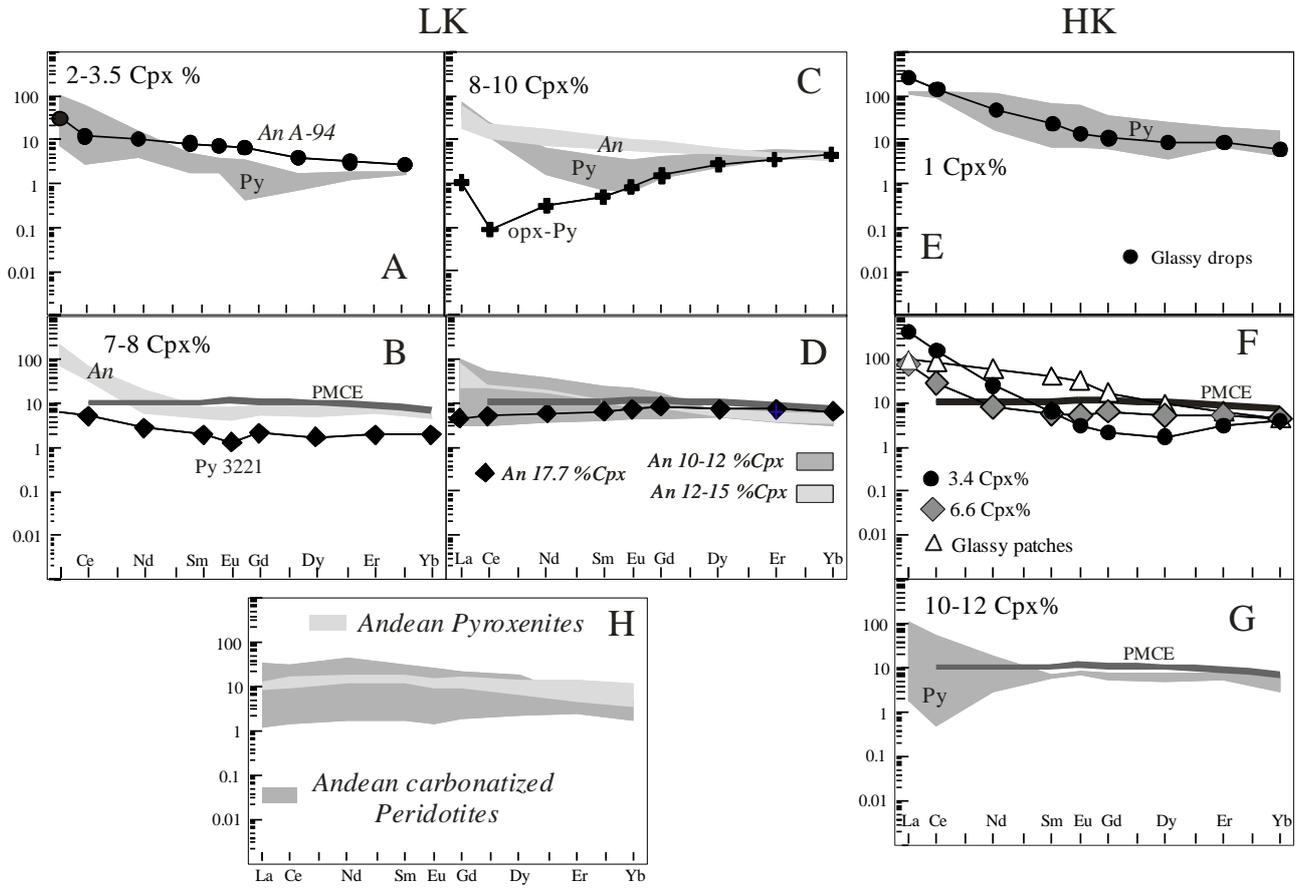


Fig. 7

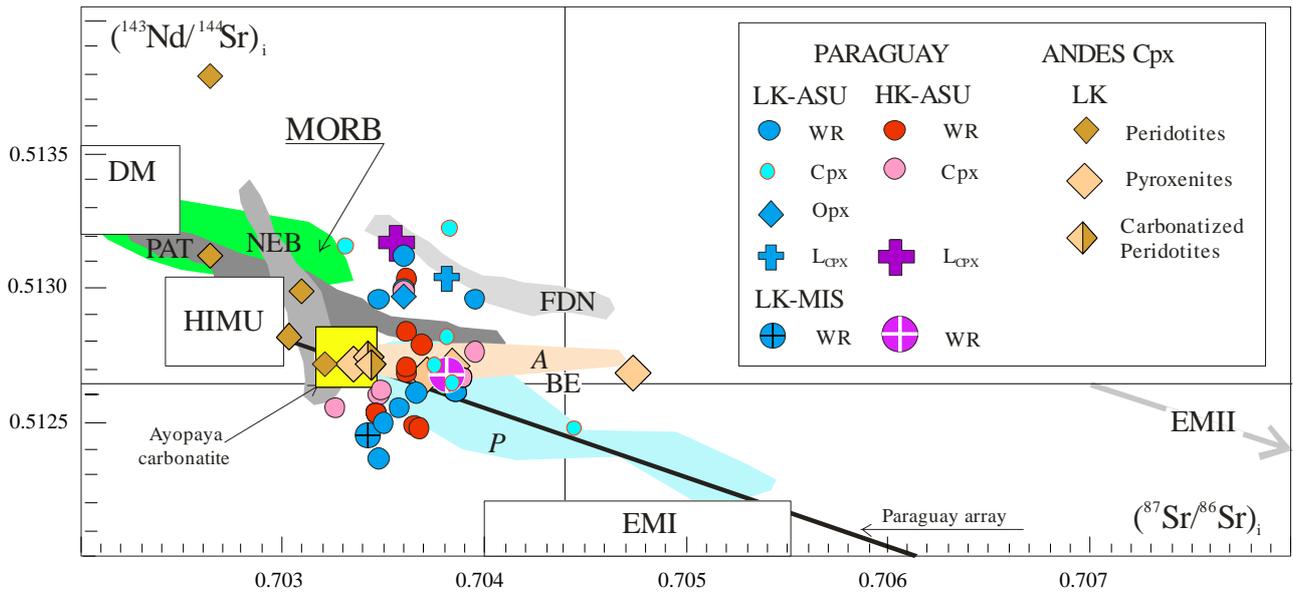


Fig. 8

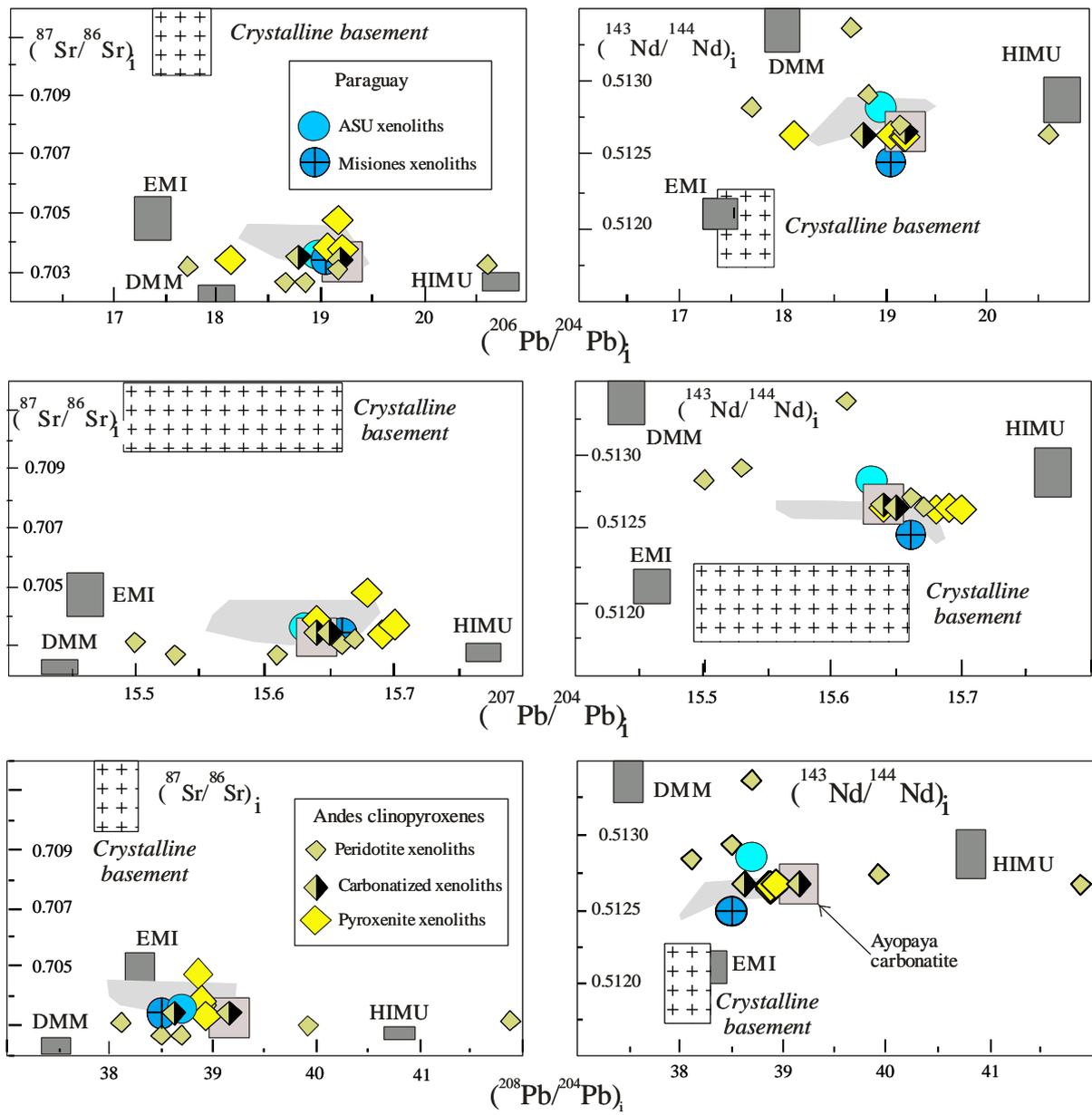


Fig. 9

Table 1

	1	2	3	4	5	6
	ASUNCIÓN	MISIONES	BELÉN	LAS CONCHAS + CADILLAL	FINCA DEL RODEO	BETANZOS
Rock-type	melanephelinite	ankaratrite	tephrite	basanite	ankaratrite	basanite
N. of Samples	31	6	3	2	3	6
wt%						
SiO ₂	42.91 (0.67)	41.43 (1.00)	44.70 (0.20)	45.25 (0.05)	41.53 (0.32)	43.35 (0.89)
TiO ₂	2.05 (0.06)	3.50 (0.25)	4.10 (0.00)	2.20 (0.00)	3.20 (0.00)	3.77 (0.23)
Al ₂ O ₃	13.41 (0.40)	13.07 (0.76)	14.63 (0.06)	14.75 (0.05)	13.60 (0.04)	13.90 (0.48)
FeO _{tot}	10.10 (0.60)	11.79 (0.29)	12.48 (0.35)	9.58 (0.05)	11.22 (0.07)	12.06 (0.49)
MnO	0.20 (0.02)	0.19 (0.01)	0.19 (0.02)	0.19 (0.00)	0.19 (0.01)	0.22 (0.03)
MgO	9.54 (0.51)	9.46 (0.58)	7.67(0.12)	10.50 (0.10)	10.97 (0.06)	8.12 (1.00)
CaO	10.45 (0.30)	11.46 (0.15)	8.40 (0.43)	10.35 (0.05)	12.33 (0.07)	10.57 (0.33)
Na ₂ O	5.80 (0.56)	4.88 (0.64)	3.67 (0.38)	4.35 (0.05)	4.27 (0.15)	4.10 (0.57)
K ₂ O	1.46 (0.39)	1.36 (0.24)	1.73 (0.21)	0.80 (0.00)	0.73 (0.12)	1.63 (0.23)
P ₂ O ₅	1.18 (0.06)	0.74 (0.02)	1.03 (0.06)	0.93 (0.01)	0.78 (0.02)	0.97 (0.15)
LOI	2.19 (0.98)	1.84 (0.03)	–	–	–	–
Sum	98.89	99.72	98.60	99.77	98.82	98.69
Notional	60	119	130	100	96	80
Age (Ma)						
mg#	0.67	0.63	0.57	0.70	5.85	2.52
Na ₂ O/ K ₂ O	3.97	3.59	2.12	5.44	0.67	0.59
ppm						
Cr	648 (65)	275 (34)	117 (30)	333 (13)	230 (10)	177 (99)
Ni	273 (15)	122 (6)	66 (12)	232 (6)	146 (2)	129 (77)
Rb	59 (9)	87.5 (11.9)	42 (16)	5.8 (2.3)	14.7 (29)	16.7 (5.4)
Ba	980 (87)	891 (62)	981 (233)	766 (15)	949 (31)	954 (76)
Th	11.0 (1.1)	7.11 (1.93)	7.88 (0.03)	5.32 (0.45)	7.78 (0.12)	9.56 (1.60)
U	2.4 (0.6)	1.47 (0.52)	2.12 (0.11)	1.45 (0.03)	1.58 (0.13)	2.73 (0.58)
Pb	10.3 (1.6)	7.31 (2.52)	7.26 (0.62)	5.55 (0.04)	7.06 (0.31)	5.47 (1.59)
Ta	7.1 (0.9)	4.0 (0.19)	–	–	–	–
Nb	101 (7)	74 (2)	84 (1)	93 (1)	71 (0)	125 (17)
Sr	1013 (98)	1025 (70)	1097 (76)	798 (4)	1080 (174)	1147 (135)
Hf	5.5 (0.8)	6.24 (0.17)	9.36 (0.09)	5.94 (0.50)	6.58 (0.14)	8.60 (0.99)
Zr	234 (41)	270 (13)	321 (1)	256 (4.24)	190 (8)	341 (59)
Y	33 (3)	27 (4)	31.1 (1.5)	20.6 (0.6)	26.7 (0.4)	26.7 (2.8)
La	119 (7)	69.5 (13.9)	68.2 (0.6)	29.1 (0.6)	63.4 (1.1)	68.5 (22.8)
Ce	186 (11)	127 (26)	135.8 (2.4)	61.6 (1.1)	117.3 (2.8)	133.6 (37.5)
Pr	20.5 (3.2)	14.0 (2.9)	15.6 (0.4)	7.6 (0.0)	13.1 (0.3)	15.3 (3.7)
Nd	63.7 (5.1)	61.5 (7.2)	60.7 (1.8)	31.5 (0.1)	50.3 (0.7)	58.7 (11.8)
Sm	11.23 (2.98)	11.01 (1.44)	11.3 (0.5)	6.5 (0.1)	9.4 (0.3)	10.6 (1.5)
Eu	2.15 (0.57)	2.37 (0.07)	3.53 (0.12)	2.10 (0.00)	2.97 (0.06)	3.35 (0.22)
Gd	5.16 (1.37)	5.06 (0.66)	9.70 (0.52)	5.85 (0.21)	8.23 (0.15)	8.78 (1.07)
Tb	0.75 (0.20)	0.93 (0.15)	1.29 (0.03)	0.84 (0.01)	1.09 (0.01)	1.16 (0.12)
Dy	4.71 (1.67)	3.61 (0.47)	6.92 (0.29)	4.68 (0.04)	5.87 (0.05)	6.29 (0.58)
Ho	–	–	1.19 (0.06)	0.86 (0.01)	1.02 (0.02)	1.10 (0.11)
Er	2.75 (0.61)	2.30 (0.13)	3.13 (0.10)	2.25 (0.06)	2.60 (0.04)	2.81 (0.25)
Tm	–	–	0.40 (0.00)	0.30 (0.01)	0.32 (0.01)	0.35 (0.02)
Yb	1.79 (0.28)	1.78 (0.10)	2.42 (0.05)	1.97 (0.01)	1.91 (0.03)	2.17 (0.11)
Lu	0.30 (0.04)	0.25 (0.02)	0.35 (0.01)	0.30 (0.01)	0.26 (0.01)	0.31 (0.01)
(Eu/Eu)*	0.75	0.85	1.01	1.03	1.01	1.03

Table 2.

Location	Asunción	Misiones	Belén	Las Conchas and Cadillal	Finca del Rodeo	Betanzos	Ayopaya carbonatite
N. Samples	10	6	3	2	3	6	1
Notional Age	60	119	130	100	96	80	98
Rb	56±20	87.5±11.9	42±16	5.8±2.3	14.7±2.9	16.7±5.4	<0.1
Sr	1080±76	1025±70	1097±76	798±4	1080±174	1147±135	13440
(⁸⁷ Sr/ ⁸⁶ Sr) _i	0.70367	0.70425	0.70380	0.70363	0.70418	0.70339	0.70329
±	0.00011	0.00016	0.00022	0.00004	0.00041	0.00022	0.00001
Sm	10.11±1.66	11.01±1.44	11.3±0.5	6.5±0.1	9.4±0.3	10.6±1.5	86.0
Nd	59.93±8.78	61.5±7.2	60.7±1.8	31.5±0.1	50.3±0.7	58.7±11.8	538.4
(¹⁴³ Nd/ ¹⁴⁴ Nd) _i	0.51268	0.51242	0.51267	0.51275	0.51274	0.51274	0.51273
±	0.00006	0.00001	0.00001	0.00001	0.00001	0.00001	0.00001
ε Sr	-13.8	-4.5	-10.7	-13.5	-5.7	-17.2	-18.4
ε Nd	2.3	-1.2	3.8	4.7	4.3	4.3	4.3
U	2.00±0.03	1.47±0.52	2.12±0.11	1.45±0.01	1.58±0.13	2.73±0.58	3.48
Th	10.50±0.63	7.11±1.93	7.88±0.003	5.32±0.45	7.78±0.12	9.56±1.60	13.15
Pb	10.06±0.26	7.31±2.52	7.26±0.62	5.55±0.04	7.06±0.31	5.47±1.59	22.16
(²⁰⁶ Pb/ ²⁰⁴ Pb) _i	18.89±0.07	18.31±0.14	19.29±0.08	18.63±0.00	19.13±0.01	19.39±0.17	19.20
(²⁰⁷ Pb/ ²⁰⁴ Pb) _i	15.67±0.01	15.68±0.00	15.57±0.01	15.56±0.00	15.58±0.01	15.59±0.01	15.64
(²⁰⁸ Pb/ ²⁰⁴ Pb) _i	38.37±0.17	37.97±0.01	39.09±0.07	38.25±0.05	39.16±0.02	39.06±0.18	39.15
μ	12.8	12.8	19.1	16.8	14.6	25.5	10.2
κ	5.4	5.0	3.8	5.6	3.8	3.6	3.9

Tab. 3.

Suite	LK		LK		LK		LK		LK		LK		LK		
Cpx% Samples	2-3%		7-8%		8-10%		10-12%		10-12%		Pyroxenite		Pyroxenite		
	P-3269; P-3227; P-M569X	A-94	P-3221	A-101; A-106; A-00179; A-6-18	P-3252; P-3254	P-3254	P-3254	A-4-147; A-00177; A-6-178b	A-98d; A-6-179; A-7-169A; A-00175	A-4-233; A-4-240a; A-4-242; A-7-169b					
ppm	Max	min		Max	min	Max	min	Opx	Max	min	Max	min	Max	min	
Rb	1.70	0.74	0.53	1.20	1.09	0.32	0.08	0.01			0.48	0.44	0.57	0.43	
Ba	0.90	0.51	1.97	0.60	28.6	7.3	0.59	0.85	0.39	3.79	2.06	5.46	1.38	29.8	10.3
Nb	0.18	0.10	2.06	0.10	12.7	0.88	0.11	0.10	0.09	1.49	0.88	3.05	1.49		
K	12.50	8.50		118				142	90	11.4					
La	33.32	0.41	4.58	23.70	66.1	1.89	22.84	10.87	0.32	18.7	5.55	29.9	1.01	3.99	2.46
Ce	47.48	2.18	9.63	23.64	58.2	4.1	20.98	8.72	0.07	19	8.65	46.3	2.51	13.20	7.66
Sr	234	38	97.6	425	295	7.79	132	106	1.70	172	93.5	245	29.4	71.4	44.8
Nd	9.30	2.21	5.98	3.36	10.9	1.73	3.85	0.88	0.18	10.2	4.78	22.3	2.15	11.3	6.84
Zr	11.50	5.07	16	0.80	52	7.3	0.53	0.13	0.16	35.9	17	112	18.4	78.4	50.2
Sm	0.92	0.31	1.64	0.85	1.66	0.37	0.80	0.14	0.09	2.33	1.10	5.07	0.83	3.45	2.10
Eu	0.28	0.12	0.51	0.29	0.62	0.081	0.27	0.05	0.06	0.79	0.45	1.59	0.32	1.13	0.70
Ti	420	160		1065				750	260	314					
Gd	0.95	0.10	1.71	1.29	2.43	0.55	1.08	0.37	0.36	2.43	1.31	4.64	1.21	4.20	2.45
Dy	0.57	0.22	1.29	1.67	2.79	0.56	1.75	0.75	0.85	2.15	1.46	2.06	1.46	4.92	2.04
Y	11	3.20	5.9	2.30	15.1	3.27	12.0	6.20	7.50	10.8	6.68	13.4	6.83	17.50	8.44
Er	0.38	0.25	0.64	1.15	1.76	0.397	1.24	0.73	0.71	1.13	0.85	1.64	0.721	2.94	0.953
Yb	0.41	0.30	0.54	0.96	1.44	0.41	1.26	0.73	0.93	1.06	0.69	1.4	0.69	2.41	0.674
PM															
Sr/Sr*	16.81	1.14	0.94	2.55	0.68	0.21	0.85	1.81	0.94	0.90	1.05	0.55	0.93	0.42	0.45
Zr/Zr*	0.23	0.39	0.35	0.03	0.79	0.62	0.02	0.02	0.06	0.61	0.51	0.72	0.89	0.87	0.97
Ti/Ti*	0.20	0.27		0.42			0.34	0.46	0.50						
Y/Y*	3.09	1.71	0.84	0.22	0.89	0.91	1.07	1.07	1.25	0.90	0.78	0.95	0.86	0.67	0.78
CH															
(La/Sm)N	22.78	1.03	1.76	17.53	25.05	3.21	17.96	48.83	2.24	5.05	3.17	3.71	0.77	0.73	0.74
(Sm/Dy)N	2.67	2.33	5.86	0.84	0.98	5.57	0.75	0.31	0.17	1.79	1.24	4.06	0.94	1.16	1.42
(Gd/Yb)N	1.87	0.27	2.56	1.08	1.36	1.08	0.69	0.41	0.31	1.85	1.53	2.67	1.42	1.41	2.93
(La/Yb)N	54.79	0.92	5.72	16.6	27.25	3.11	12.17	10.04	0.23	11.89	5.42	14.40	0.99	1.12	2.46
(Eu/Eu)*	0.91	1.65	0.92	0.85	0.94	0.62	0.89	0.63	0.89	1.01	1.14	0.98	0.97	0.91	0.94

Tab. 3. (cont.)

Suite	LK		LK	LK		HK		HK	HK	HK	HK	HK	HK
Cpx%	12-15%		17.7%	Carbonatized Peridotite		1%		3.4%	6.6%	10-12%	Glassy drops		Glassy patches
Samples	A-102; A-7-168b		A-104	A-57; A-113-c; A-4-229; A-4-301a		P-3284 P-SI-2A		P-3211	P-3307	P-3311; P-3222		P-3284	P-3211
ppm	Max	min		Max	min	Max	min			Max	min		
Rb	0.24	0.20	0.28	2.65	0.45	1.40	1.10	1.80	0.30	1.30	0.07	120	87
Ba	3.41	2.53	1.03	37.5	1.93	4.47	2.10	3.50	0.25	0.60	0.50	1781	823
Nb	8.6	6.4			0.49	0.30	0.25	0.70	0.30	0.20	0.11	155	277
K						3.81	81	397	10	90	30	82453	20091
La	28.12	7.09	1.36	10.7	0.36	38	323	119.4	23	36.14	0.58	77.39	29
Ce	20.6	19.9	3.95	25.5	1.16	98	68.4	121.5	22	45.55	0.38	106.25	68
Sr	119	110	70.1	268	132	525	248	977	130	336	18.4	1414	990
Nd	11.5	11	3.62	25.3	1.04	62	8.89	15.13	4.86	10.37	1.67	27.74	34
Zr	92.8	7.05	30.1	50.2	8.5	97	12.77	2.60	14	14.00	11.10	189	109
Sm	2.24	2.07	1.25	5.94	0.31	13.2	1.36	1.23	1.16	1.26	1.20	4.16	7.4
Eu	0.654	0.65	0.563	1.90	0.11	4.37	0.48	0.23	0.39	0.56	0.54	1.00	2.35
Ti						1482	4.83	405	1253	2157	955	4888	2338
Gd	2.21	1.88	2.12	5.35	0.46	9.71	1.56	0.56	1.60	2.02	1.45	2.92	4.70
Dy	2.29	1.46	2.40	3.31	0.67	7.8	1.14	0.55	1.54	2.58	1.63	2.61	3.18
Y	11.8	6.69	12.7	13.4	3.98	39	12.9	3.60	9.00	16.80	10.40	14.70	18.0
Er	1.38	0.75	1.54	1.19	0.49	3.95	1.36	0.69	1.12	1.63	1.10	1.68	1.29
Yb	1.24	0.67	1.35	1.02	0.33	3.24	0.90	0.82	0.96	1.38	0.60	1.23	0.98
PM													
Sr/Sr*	0.56	0.54	1.36	0.77	8.86	0.49	0.52	1.16	0.77	0.10	1.21	1.67	1.48
Zr/Zr*	0.28	1.07	0.87	0.23	0.24	0.03	0.40	0.03	0.40	0.22	0.50	1.13	0.47
Ti/Ti*				0.05	0.14	0.27	0.38	0.27	0.38	0.49	0.26	0.69	0.16
Y/Y*	0.85	0.90	0.60	0.91	2.19	0.72	0.90	0.72	0.90	1.04	1.02	0.92	1.13
CH													
(La/Sm)N	1.13	0.73	0.73	1.81	14.95	61.01	12.47	61.01	12.47	18.05	0.30	11.72	2.47
(Sm/Dy)N	2.96	0.76	1.16	2.79	1.97	3.69	1.24	3.69	1.24	0.81	1.22	2.63	3.84
(Gd/Yb)N	4.23	1.13	1.41	2.43	1.41	0.55	1.39	0.55	1.39	1.18	1.95	1.91	3.87
(La/Yb)N	7.07	7.13	1.12	7.07	0.73	7.91	24.17	98.21	16.67	17.66	0.65	42.44	19.95
(Eu/Eu)*	1.01	0.89	0.91	1.13	1.01	0.74	0.88	0.74	0.88	1.07	1.25	0.83	1.14

Tab. 4.

Suite/Sample	Cpx%	Rb	Sr	Sm	Nd	(⁸⁷ Sr/ ⁸⁶ Sr) _m	(¹⁴³ Nd/ ¹⁴⁴ Nd) _m	(⁸⁷ Sr/ ⁸⁶ Sr) _i	(¹⁴³ Nd/ ¹⁴⁴ Nd) _i	TDM Ma
LK										
MIS-WR										
PS-558B-MS	1.0	1.70	9.77	5.550	32.18	0.70438	0.51253	0.70343	0.51245	822
ASU-WR										
3269	1.9	1.74	9.94	0.010	0.29	0.70452	0.51267	0.70360	0.51265	371
3227	3.2	1.01	8.75	0.054	0.55	0.70446	0.51301	0.70385	0.51296	135
3253	3.5	0.65	8.98	0.036	0.20	0.70399	0.51305	0.70395	0.51296	140
3213	4.0	1.20	14.50	0.038	0.22	0.70392	0.51246	0.70348	0.51237	921
3199	4.5	1.92	16.94	0.012	0.20	0.70426	0.51264	0.70366	0.51261	427
3221	7.1	1.99	19.90	0.042	0.31	0.70410	0.51263	0.70357	0.51256	581
3252	8.9	0.70	6.12	0.040	0.150	0.70420	0.51311	0.70360	0.51300	109
3254	8.9	0.48	2.11	0.027	0.040	0.70480	0.51346	0.70360	0.51312	250
3192	9.8	1.07	6.80	0.040	0.25	0.70433	0.51258	0.70350	0.51250	713
ASU-OPX										
3252		-	-	0.030	0.18	0.70360	0.51305	0.70360	0.51297	131
ASU CPX										
3269		0.020	415.62	0.594	6.64	0.70444	0.51253	0.70444	0.51248	483
3227		0.022	199.43	0.469	4.80	0.70375	0.51277	0.70375	0.51272	364
3221		0.001	12.74	0.575	4.21	0.70384	0.51272	0.70384	0.51265	403
3254		1.45	9.80	0.089	0.18	0.70461	0.51348	0.70383	0.51322	627
3252		0.011	103.30	0.650	3.70	0.70382	0.51291	0.70382	0.51282	327
3192		0.136	45.85	2.595	4.71	0.70332	0.51344	0.70331	0.51316	386
*Leaching sol						0.70382	0.51304			
ANDES-CPX										
A-106	7.1	1.94	7.79	0.371	1.73	0.70368	0.51387	0.70264	0.51379	-
6-181	7.1	1.09	224	0.973	10.9	0.70323	0.51275	0.70321	0.51272	370
4-147-m	8.6	0.44	110	1.100	4.78	0.70304	0.51290	0.70303	0.51282	477
A-104	17.7	0.28	70.1	1.25	3.62	0.70312	0.51311	0.70310	0.51298	591
6-180a	26.1	0.68	39.2	0.948	2.85	0.70271	0.51325	0.70264	0.51312	-
CARB-L-CPX										
A-113e		0.45	223	2.69	13.70	0.70345	0.51280	0.70344	0.51272	535
4-301a		2.65	268	5.94	25.30	0.70346	0.51283	0.70342	0.51274	639
PYROX-CPX										
4-233		0.43	44.8	2.10	6.84	0.70339	0.51284	0.70335	0.51272	1411
4-240a		0.51	71.4	2.56	8.55	0.70374	0.51280	0.70371	0.51268	1496
4-242		0.57	62.7	2.18	7.35	0.70478	0.51281	0.70474	0.51269	1313
7-169b		0.49	54.9	3.45	11.30	0.70388	0.51283	0.70383	0.51271	1430
HK										
MIS-WR										
SI-2-A	1.0	4.25	40.77	1.36	8.89	0.70436	0.51276	0.70381	0.51268	473
ASU-WR										
3284	1.0	6.58	63.12	0.390	2.65	0.70401	0.51261	0.70346	0.51254	636
Leaching sol		120	1414	11.16	27.74	0.70401	0.51325	0.70380	0.51315	619
3220	1.1	0.59	14.22	0.076	0.56	0.70383	0.51276	0.70361	0.51269	436
3301	2.1	4.23	22.97	0.140	1.13	0.70464	0.51254	0.70367	0.51248	648
3313	2.4	3.93	58.02	0.210	2.42	0.70397	0.51288	0.70361	0.51284	249
3211	3.4	2.85	44.15	0.180	2.16	0.70380	0.51258	0.70346	0.51254	518
3307	6.6	0.05	28.20	0.190	2.05	0.70366	0.51254	0.70365	0.51249	573
3311	8.8	3.86	68.15	0.320	2.2	0.70399	0.51286	0.70369	0.51279	340
3222	10.9	1.38	14.33	0.470	1.12	0.70412	0.51325	0.70361	0.51304	434
3288	11.9	4.62	41.22	0.150	0.90	0.70420	0.51279	0.70361	0.51271	467
ASU CPX										
3284		1.73	653.79	16.145	84.34	0.70350	0.51272	0.70349	0.51262	638
3211		0.11	977.26	1.993	26.04	0.70326	0.51260	0.70326	0.51256	488
3307		0.04	173.74	0.841	4.13	0.70389	0.51277	0.70389	0.51267	607
3311		15.96	309.91	0.980	6.70	0.70422	0.51283	0.70395	0.51276	376
3222		0.12	44.8	0.873	2.40	0.70361	0.51317	0.70360	0.51299	-
3288		28.66	250.51	2.410	16.80	0.70408	0.51267	0.70348	0.51260	556
Ayopaya										
140		<0.1	13440	86.0	538.4	0.70329	0.51279	0.70329	0.51271	445

Table 5.

CPX	U	Th	Pb	$(^{206}\text{Pb}/^{204}\text{Pb})_m$	$(^{207}\text{Pb}/^{204}\text{Pb})_m$	$(^{208}\text{Pb}/^{204}\text{Pb})_m$	μ	κ	$(^{206}\text{Pb}/^{204}\text{Pb})_i$	$(^{207}\text{Pb}/^{204}\text{Pb})_i$	$(^{208}\text{Pb}/^{204}\text{Pb})_i$
A-104	0.028	0.03	0.39	17.78	15.5	38.13	4.5	1.1	17.71	15.50	38.11
A-106	0.242	1.02	6.83	18.70	15.61	38.74	2.4	4.4	18.66	15.61	38.69
A-147-m	0.180	0.07	0.22	20.02	15.66	40.04	55.0	0.4	19.16	15.66	39.93
6-180-a	0.089	0.33	1.82	18.89	15.53	38.56	3.1	3.8	18.84	15.53	38.50
6-181	0.368	1.40	3.42	20.73	15.68	42.03	7.4	3.9	20.61	15.67	41.89
CPX in carbonatized peridotites											
A-113e	0.087	0.05	0.95	18.87	15.65	38.79	6.0	0.6	18.78	15.65	38.62
4-301a	0.219	1.2	10.8	19.21	15.64	39.20	1.3	5.7	19.19	15.64	39.16
CPX in Pyroxenite											
4-233	0.074	0.29	5.9	18.13	15.65	38.96	0.7	4.5	18.12	15.69	38.94
4-240a	0.119	0.52	2.32	19.26	15.70	38.96	3.3	4.5	19.21	15.70	38.88
4-242	0.161	0.35	3.56	19.22	15.68	38.89	2.9	2.3	19.17	15.68	38.85
7-169b	0.066	0.26	1.34	19.11	15.64	38.94	3.2	4.1	19.06	15.64	38.88
ASU Xenolith (LK)											
3252				18.96	15.63	38.7					
MISIONES Xenolith (LK)											
PS-558B-X				19.05	15.66	38.50					